

Extended Higgs Sector for Neutrino Mass, Baryogenesis and Dark Matter



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S.K., N. Machida, T. Shindou, arXiv:1405.5834

Multi-Higgs Models, Lisbon, 2-5. Sep. 2014

Phenomena beyond the SM

We already know **BSM** phenomena:

- Neutrino oscillation

$$\Delta m^2 \sim 8 \times 10^{-5} \text{ eV}^2, \Delta m^2 \sim 2 \times 10^{-3} \text{ eV}^2$$

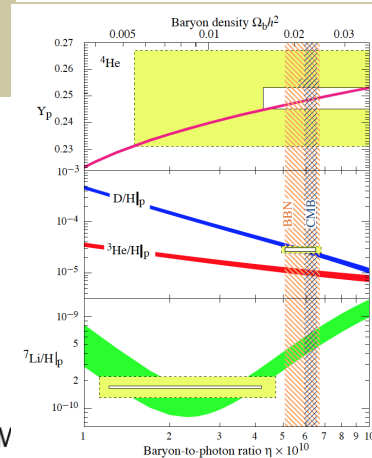
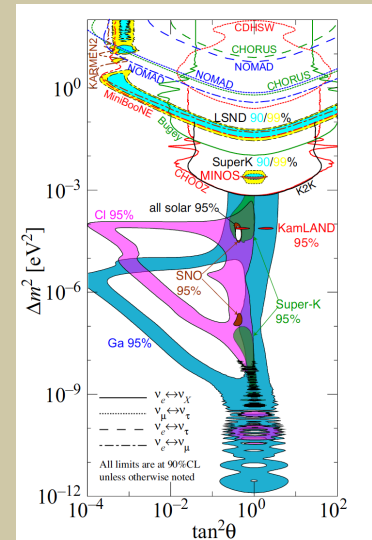
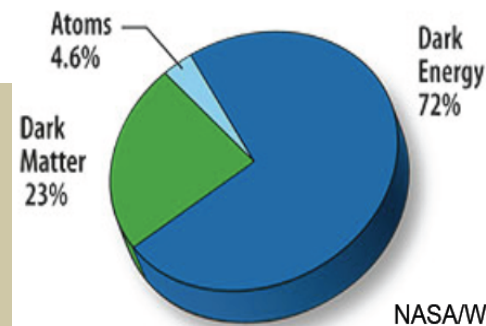
- Dark Matter

$$\Omega_{\text{DM}} h^2 \sim 0.12$$

- Baryon Asymmetry of the Universe

$$n_B/n_\gamma \sim 6 \times 10^{-10}$$

New physics is necessary!
Which scale?



If NP appears at the **TeV scale**, it should have a strong connection with the physics behind the **Higgs sector**

$$\eta_B = \frac{n_B}{n_\gamma} = \frac{n_b - n_{\bar{b}}}{n_\gamma}$$

Dynamics behind the 125 GeV Higgs

- **Weak and Light Scenario**

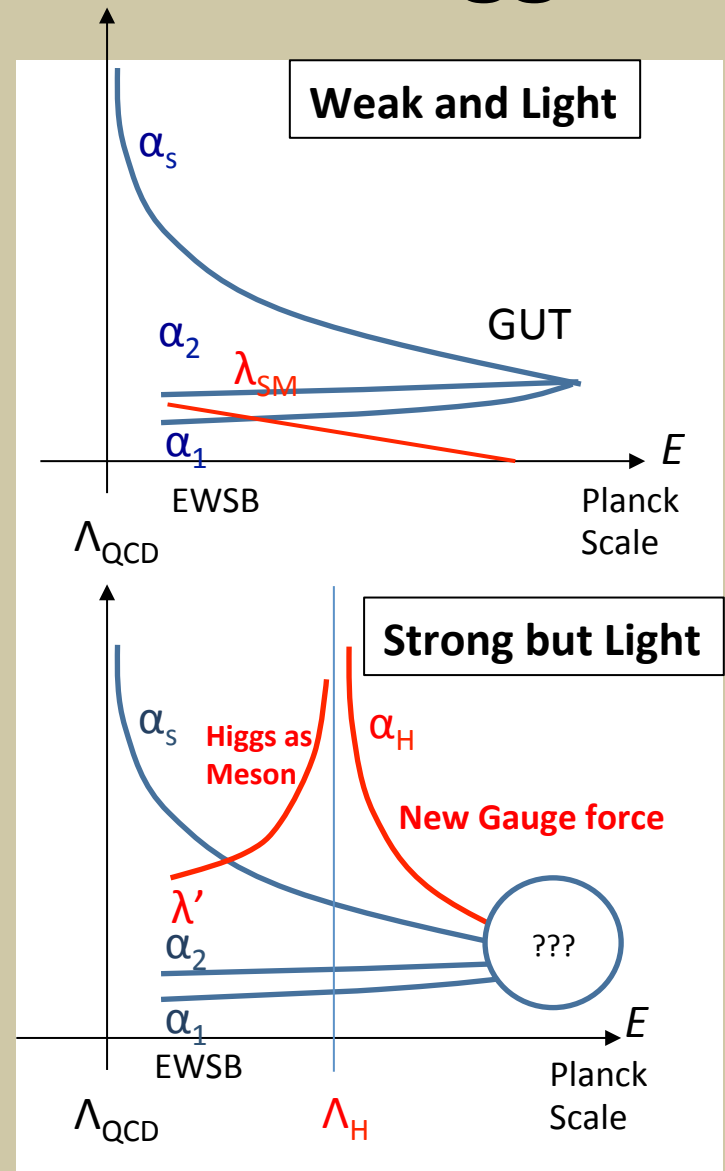
- Perturbative
- Grand Desert
- Traditional Grand Unification

$$m_h^2 \propto \lambda v^2$$

- **Strong but Light Scenario**

- IR theory:
Higgs as a **composite** field
Landau pole at Λ_H
- UV theory:
A new gauge symmetry with confinement at Λ_H

$$m_h^2 \propto \frac{\lambda'^2}{(4\pi)^2} v^2$$



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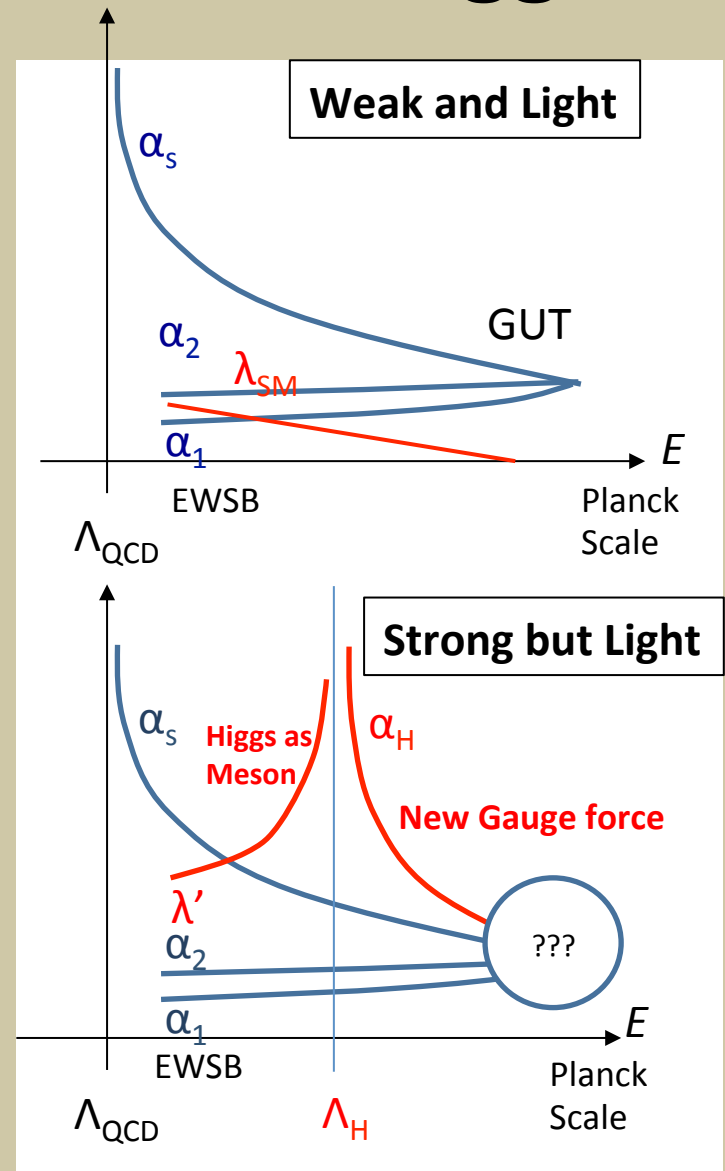
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Electroweak Baryogenesis

Sakharov's conditions:

B Violation

C and CP Violation

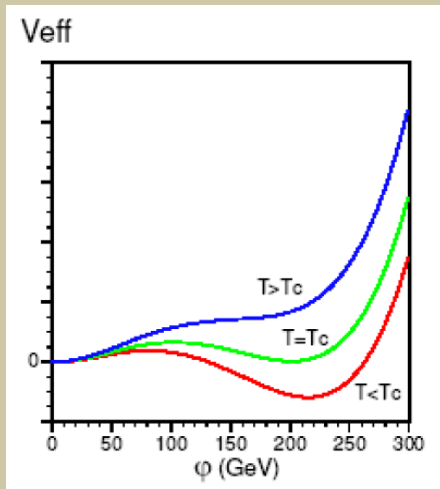
Departure from Equilibrium

→ **Sphaleron transition at high T**

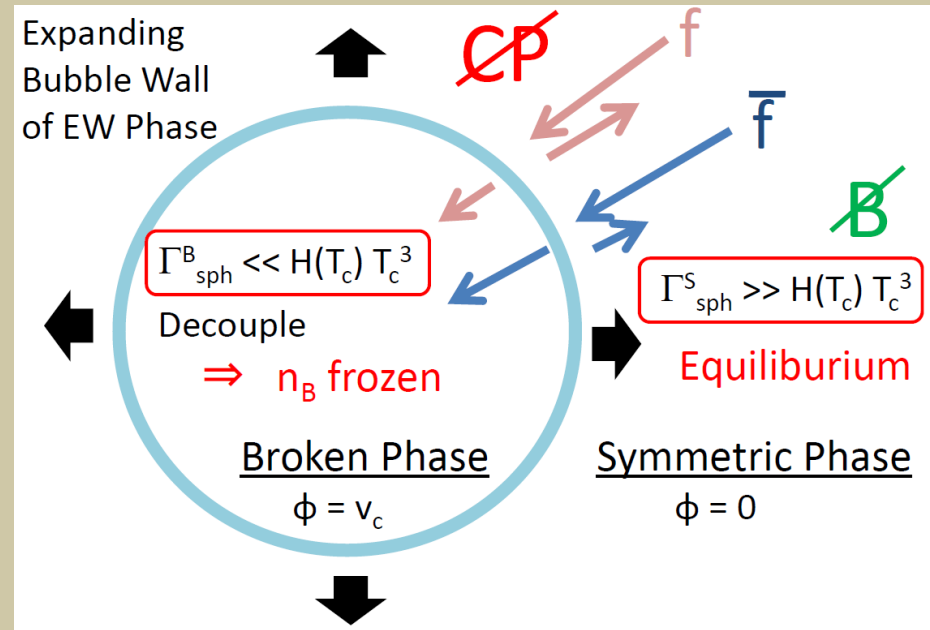
→ **CP Phases in extended scalar sector**

→ **1st Order EW Phase Transition**

$$\begin{aligned}\Gamma &\sim e^{-E_{\text{sph}}/T} \quad (T < T_c) \\ \Gamma &\sim \kappa(\alpha_W T)^4 \quad (T_c < T)\end{aligned}$$



Quick sphaleron decoupling is required to retain sufficient baryon number in Broken Phase



(Sphaleron Rate) < (Expansion Rate)

$$\frac{\varphi_c}{T_c} \gtrsim 1$$

Condition of Strong 1st OPT ($\phi_c/T_c > 1$)

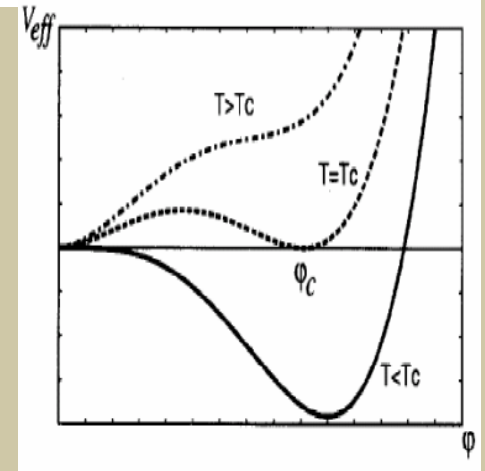
Finite Temperature Potential

$$V_T(\phi, T) = D(T^2 - T_0^2)\phi^2 - ET\phi^3 + \frac{\lambda_T}{4}\phi^4 + \dots$$

$$\phi_c/T_c > 1 \Rightarrow 2E/\lambda_{T_c} > 1$$

EWBG not viable in the SM with 125GeV Higgs

$$E = \frac{1}{12\pi v^3}(6m_W^3 + 3m_Z^3) \Rightarrow \mathbf{m_h < 80 \text{ GeV!}}$$



Condition of Strong 1st OPT ($\phi_c/T_c > 1$)

Finite Temperature Potential

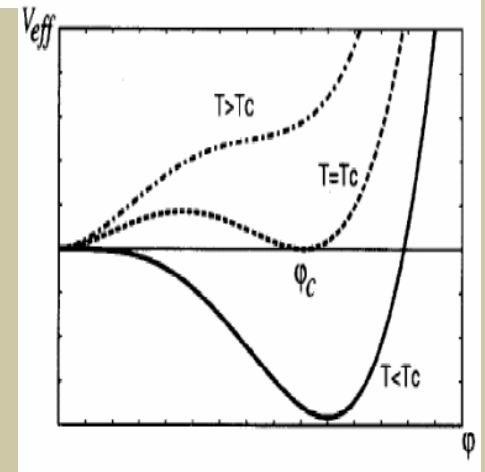
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Muti-Higgs models can satisfy the condition



$$V(\phi_{SM}, \Phi) = \dots + \lambda_{SM} \phi_{SM}^4 + \lambda' \phi_{SM}^2 \Phi^2 + \dots$$

$$E \propto \frac{m_\Phi^3}{\lambda_T} = \frac{(\lambda')^{3/2}}{\lambda_T} \Rightarrow \lambda' \gg \lambda_{SM} = 0.13 \text{ (for } m_h = 126 \text{ GeV)}$$

$$\phi_c/T_c > 1 \text{ when } \lambda' > O(1)$$

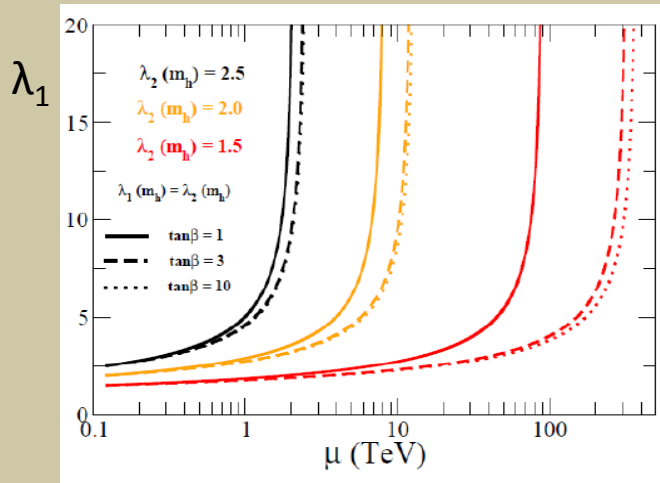
Strong-but-Light !

EW Phase Transition and Landau Pole

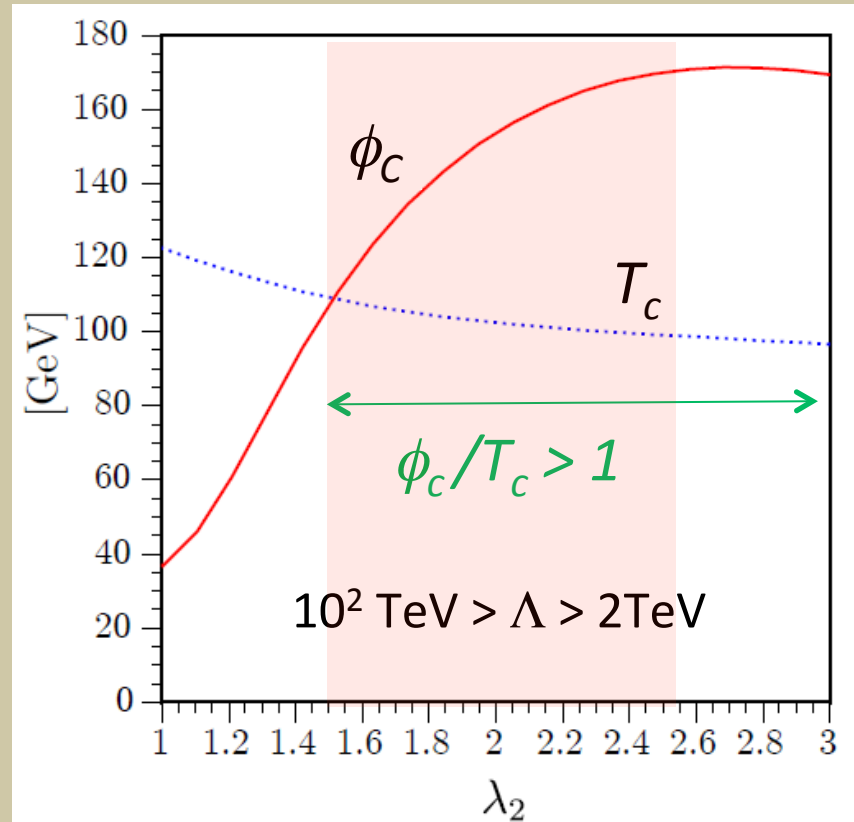
Strong 1st OPT $\rightarrow$ large λ' at EW
 $\rightarrow$ **Landau pole**

Ex) SUSY 4HDM+ Ω

$$W = \lambda_1 H_u H_u' \Omega_1 + \lambda_2 H_d H_d' \Omega_2$$



S.K., E. Senaha, T. Shindou 2011



$$\phi_c/T_c > 1 \Rightarrow \Lambda_{\text{cutoff}} < 100 \text{ TeV}$$

What is the fundamental theory above the Landau Pole?

Neutrino Mass

Neutrino Mass Term (= Effective dim-5 operator)

$$\mathcal{L}^{\text{eff}} = (c_{ij}/M) \nu_L^i \nu_L^j \phi \phi$$

$$\langle \phi \rangle = v = 246 \text{ GeV}$$

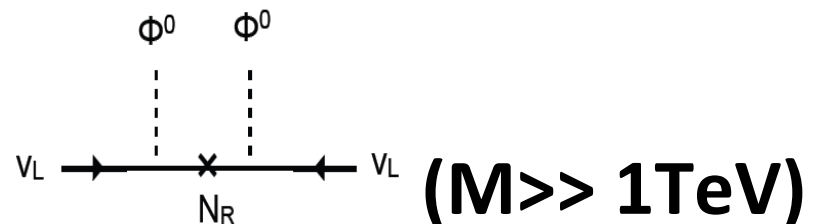
Mechanism for tiny masses:

$$m_{ij}^\nu = (c_{ij}/M) v^2 < 0.1 \text{ eV}$$

Seesaw (tree level)

$$m_{ij}^\nu = y_i y_j v^2 / M$$

Minkowski
Yanagida
Gell-Mann et al



Quantum Effects (Radiative seesaw)

N-th order of perturbation

$$m_{ij}^\nu = [g^2/(16\pi^2)]^N c_{ij} v^2 / M$$

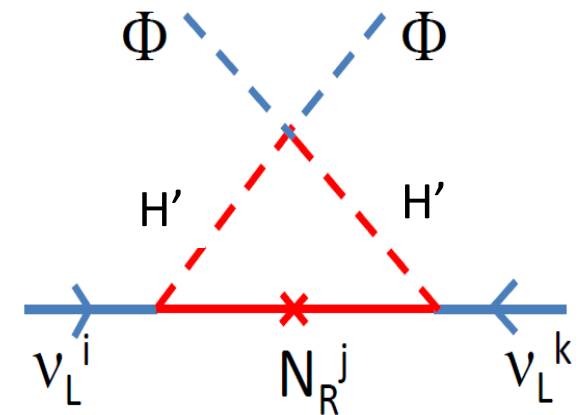
(M can be 1 TeV)

Radiative seesaw with Z_2

Z_2 -parity plays roles: 1. **No tree-level Yukawa** (Radiative neutrino mass)
2. **Stability** of the lightest Z_2 -odd particle (WIMP)

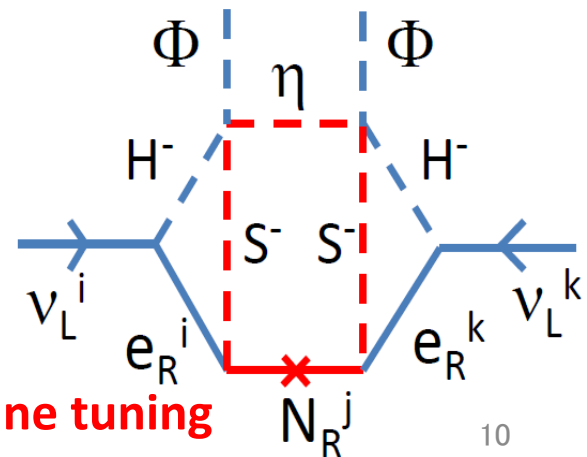
Ex1) 1-loop *Ma (2006)*

- Simplest model
- SM + **Inert scalar doublet (H')** + N_R
- DM candidate [H' or N_R]



Ex2) 3-loop *Aoki-SK-Seto (2008)*

- Neutrino mass from **$O(1)$** coupling
- 2 Higgs doublets + $\eta^0 + S^+ + N_R$
- DM candidate [η^0 (or N_R)]
- Electroweak Baryogenesis



All 3 problems may be solved by TeV physics w/o fine tuning

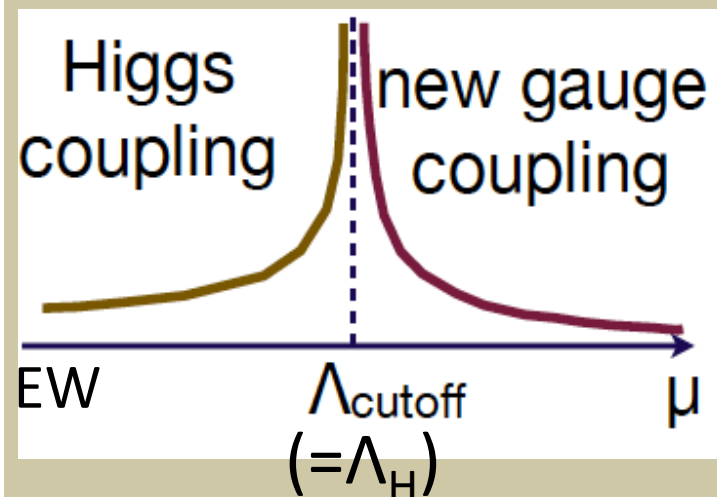
Problems in these TeV models

- Radiative Seesaw models
 - We may be able to explain all BSM phenomena
Neutrino mass, Dark Matter, Baryon Asymmetry
 - TeV scale $\Rightarrow$ Testable at colliders (Big bonus)
- However, there are problems in such scenario
 - EW Baryogenesis $\Rightarrow$ Landau pole !
 - Many scalars introduced $\Rightarrow$ Ad hoc !
 - Non-SUSY $\Rightarrow$ Hierarchy Problem
- Can we have a consistent UV theory?

This is the subject of my talk

Outline of the Model

- Origin of the Higgs force (λ) is the **$SU(N_c)$** SUSY gauge symmetry ($N_c=2, N_f=3$)
[Similar to Minimal SUSY Fat Higgs model]
Harnik, et al
- **Confinement** ($N_f = N_c+1$) occurs at Λ_H
($\sim$ Landau Pole) *Intriligator and Seiberg*
- At low energy **4 doublets + many singlets** appear with a coupling **λ** (Higgs as Meson)
- $\lambda(\text{EW})$ is set by **EWBG $\phi_c/T_c > 1$** (strong) but within perturbative $\Rightarrow \Lambda_H = O(10) \text{ TeV}$ *SK, Shindou, Yamada 2012*



By adding the **Z_2 parity** to the model, and by introducing a **RH Neutrino**, a radiative seesaw scenario is **realized** at TeV scale

SUSY $SU(2)_H$ gauge theory

Minimal model for confinement ($N_c=2, N_f=3$)

→ 3 pairs of $SU(2)_H$ fundamental rep. T_i ($i=1-6$)

Six $SU(2)_H$ doublets T_i
charged under the SM
gauge groups and Z_2 -parity

Put current mass terms
to give masses of T_i

$$W_m = m_1 T_1 T_2 + m_3 T_3 T_4 + m_5 T_5 T_6$$

Field	$SU(2)_L$	$U(1)_Y$	Z_2
$\begin{pmatrix} T_1 \\ T_2 \end{pmatrix}$	2	0	+
T_3	1	+1/2	+
T_4	1	-1/2	+
T_5	1	+1/2	-
T_6	1	-1/2	-

SK, T. Shindou, T. Yamada, 2012

Effective Theory

- It is strongly coupled at Λ_H , and T_i ($i=1-6$) are confined
K. Intriligator and N. Seiberg (1996)
- Below Λ_H the theory is described by **15** Meson Superfields

$$M_{ij} = T_i T_j$$

- Effective Superpotential**

$$W_{eff} = \frac{1}{\Lambda^3} \epsilon_{ijklmn} M_{ij} M_{kl} M_{mn} + m_1 M_{12} + m_3 M_{34} + m_5 M_{56}$$

- Using **Naïve Dimensional Analysis (NDA)**, it is rewritten by canonically normalized fields

$$W_{eff} \simeq \lambda \epsilon_{ijklmn} \hat{M}_{ij} \hat{M}_{kl} \hat{M}_{mn} + \frac{m_1 \Lambda_H}{4\pi} \hat{M}_{12} + \frac{m_3 \Lambda_H}{4\pi} \hat{M}_{34} + \frac{m_5 \Lambda_H}{4\pi} \hat{M}_{56}$$

- The coupling λ becomes non-perturbative at Λ_H

$$\lambda(\mu = \Lambda_H) \simeq 4\pi \quad \text{Naïve Dimensional Analysis}$$

Higgs fields as Mesons

Fifteen mesons $M_{ij} = T_i T_j$ can be identified as the MSSM Higgs superfields and extra superfields

Additional Super-fields	{		Field	$SU(2)_L$	$U(1)_Y$	Z_2
		MSSM Higgs doublets	H_u	2	+1/2	+
			H_d	2	-1/2	+
		Extra Higgs doublets	Φ_u	2	+1/2	-
			Φ_d	2	-1/2	-
		Charged Higgs singlets	Ω^+	1	+1	-
			Ω^-	1	-1	-
		Z ₂ -even Higgs singlets	N, N_Φ, N_Ω	1	0	+
		Z ₂ -odd Higgs singlets	ζ, η	1	0	-

Superpotential is rewritten as

$$W_{eff} = \lambda \left\{ N(H_u H_d + v_0^2) + N_\Phi(\Phi_u \Phi_d + v_\Phi^2) + N_\Omega(\Omega^+ \Omega^- + v_\Omega^2) - N N_\Phi N_\Omega - N_\Omega \zeta \eta + \zeta H_d \Phi_u + \eta H_u \Phi_d - \Omega^+ H_d \Phi_d - \Omega^- H_u \Phi_u \right\}$$

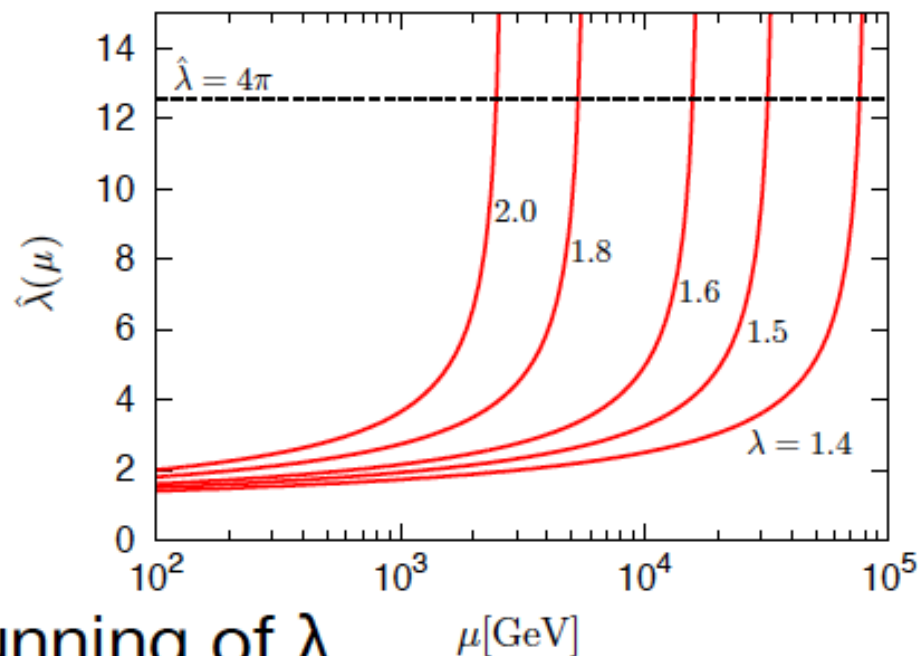
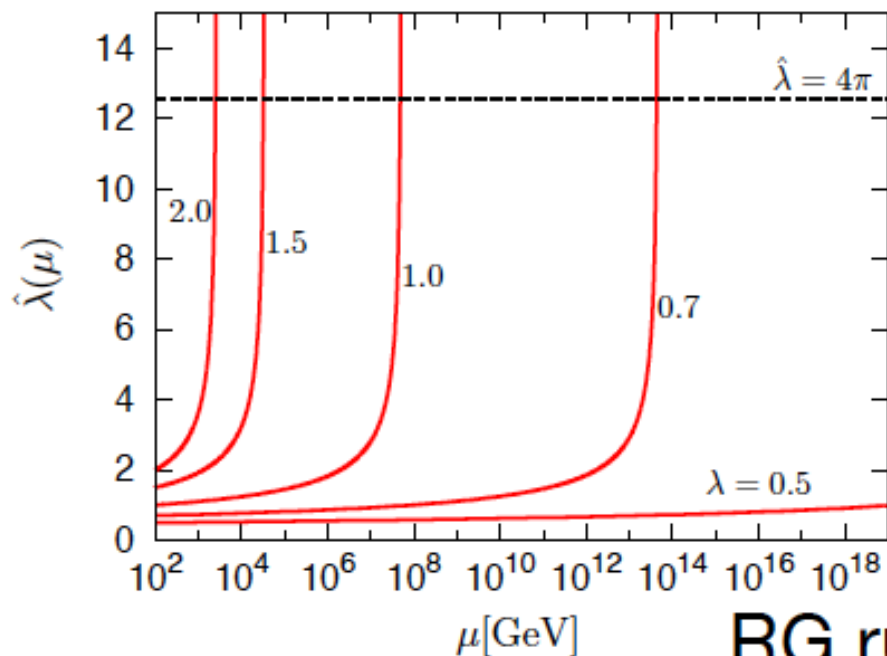
The low energy theory is **4HDM+Singlets** but with a common λ ! 15

MSSM-like Higgs doublets

$$W = -\mu H_u H_d - \mu_\Phi \Phi_u \Phi_d - \mu_\Omega (\Omega_+ \Omega_- - \zeta \eta)$$

$$+ \hat{\lambda} \{ H_d \Phi_u \zeta + H_u \Phi_d \eta - H_u \Phi_u \Omega_- - H_d \Phi_d \Omega_+ \}$$

$$\hat{\lambda}(\Lambda_H) \simeq 4\pi \text{ (Naive dimensional analysis)}$$

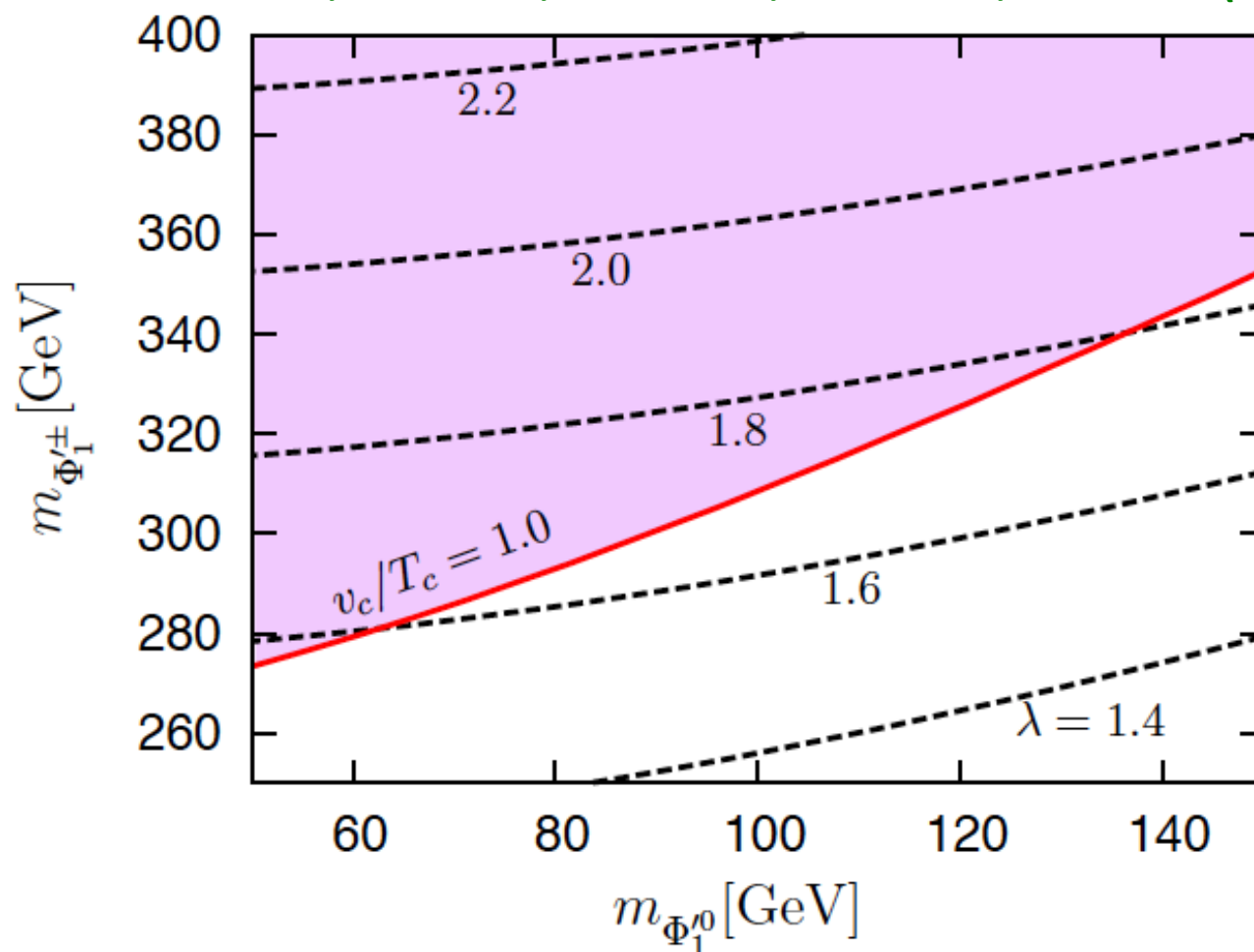


RG running of λ

$\lambda = \lambda(\mu_{EW})$ determines the cutoff scale

1st order EWPT

S.K., E. Senaha, T. Shindou, T. Yamada, JHEP1305 (2013) 066.



$$\varphi_c/T_c > 1 \implies \lambda \gtrsim 1.5 \quad (\Lambda_H \lesssim 20\text{TeV})$$

Neutrino Mass

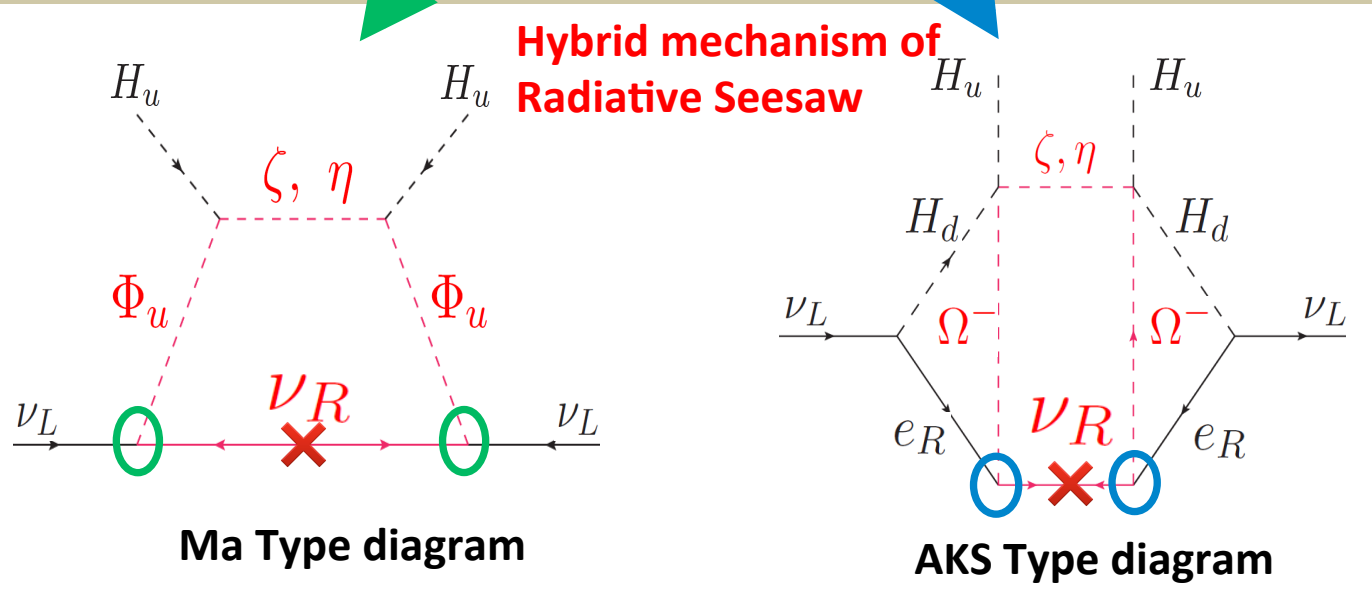
Mass generation mechanism is **hybrid** of Ma model and AKS model

Two tiny mass scales are naturally induced

One generation of ν_R is enough

All scalars in the Higgs sector contribute to the neutrino masses

$$W_{\text{eff}}^N = \frac{\kappa}{2} N \nu_R^c \nu_R^c + \underbrace{(y_N^i)}_{\text{green circle}} \nu_R^c L_i \Phi_u + \underbrace{(h_N^i)}_{\text{blue circle}} \nu_R E_i^c \Omega^- + \frac{M}{2} \nu_R^c \nu_R^c$$



$$\Delta m_{\text{sol}}^2 \sim 7 \times 10^{-5} \text{eV}^2$$

$$\Delta m_{\text{atm}}^2 \sim 2 \times 10^{-3} \text{eV}^2$$

$$\sin^2 2\theta_{12} = 0.857 \pm 0.024$$

$$\sin^2 2\theta_{23} > 0.95$$

$$\sin^2 2\theta_{13} = 0.095 \pm 0.010$$

Benchmark parameter set

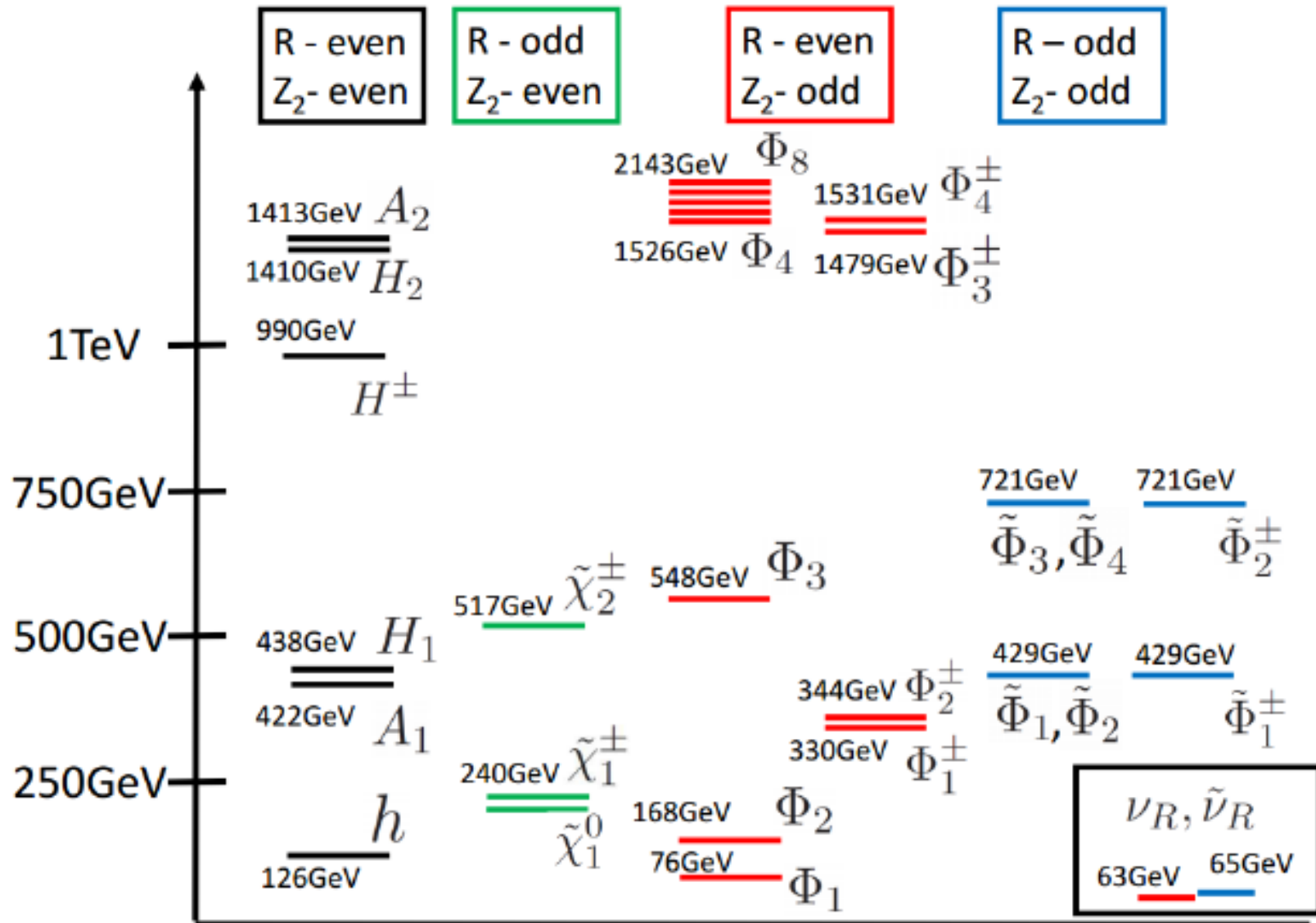
$\lambda, \tan \beta, \text{ and } \mu\text{-terms}$
$\lambda = 1.8 \ (\Lambda_H = 5 \text{ TeV}) \quad \tan \beta = 15 \quad \mu = 250 \text{ GeV} \quad \mu_\Phi = 550 \text{ GeV} \quad \mu_\Omega = -550 \text{ GeV}$
$Z_2\text{-even Higgs sector}$
$m_h = 126 \text{ GeV} \quad m_{H^\pm} = 990 \text{ GeV} \quad m_N^2 = (1050 \text{ GeV})^2 \quad A_N = 2900 \text{ GeV}$
$Z_2\text{-odd Higgs sector}$
$\bar{m}_{\Phi_u}^2 = \bar{m}_{\Omega_-}^2 = (175 \text{ GeV})^2 \quad \bar{m}_{\Phi_d}^2 = \bar{m}_{\Omega_+}^2 = \bar{m}_\zeta^2 = (1500 \text{ GeV})^2 \quad \bar{m}_\eta^2 = (2000 \text{ GeV})^2$ $B_\Phi = B_\Omega = A_\zeta = A_\eta = A_{\Omega^+} = A_{\Omega^-} = m_{\zeta\eta}^2 = 0 \quad B_\zeta^2 = (1400 \text{ GeV})^2 \quad B_\eta^2 = (700 \text{ GeV})^2$
$\text{RH neutrino and RH sneutrino sector}$
$m_{\nu_R} = 63 \text{ GeV} \quad m_{\tilde{\nu}_R} = 65 \text{ GeV} \quad \kappa = 0.9$ $y_N = (3.28i, 6.70i, 1.72i) \times 10^{-6} \quad h_N = (0, 0.227, 0.0204)$
$\text{Other SUSY SM parameters}$
$m_{\tilde{W}} = 500 \text{ GeV} \quad m_{\tilde{q}} = m_{\tilde{\ell}} = 5 \text{ TeV}$

Output

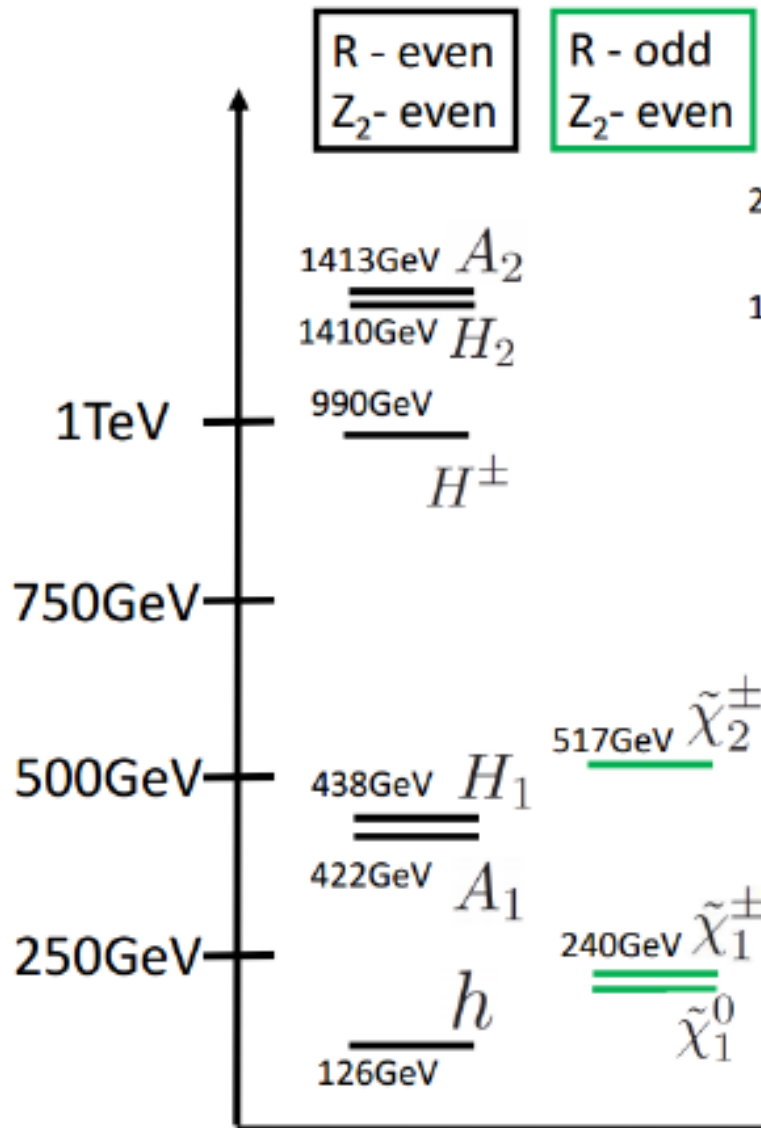
Non-decoupling effects
$\varphi_c/T_c = 1.3 \quad \lambda_{hhh}/\lambda_{hhh} _{\text{SM}} = 1.2 \quad \text{B}(h \rightarrow \gamma\gamma)/\text{B}(h \rightarrow \gamma\gamma) _{\text{SM}} = 0.78$
Neutrino masses and the mixing angles
$(m_1, m_2, m_3) = (0, 0.0084 \text{ eV}, 0.0050 \text{ eV}) \quad \sin^2 \theta_{12} = 0.32 \quad \sin^2 \theta_{23} = 0.50 \quad \sin \theta_{13} = 0.14$
LFV processes
$\text{B}(\mu \rightarrow e\gamma) = 3.6 \times 10^{-13} \quad \text{B}(\mu \rightarrow eee) = 5.6 \times 10^{-16}$
Relic abundance of the DM
$\Omega_{\nu_R} h^2 = 0.055 \quad \Omega_{\bar{\nu}_R} h^2 = 0.065 \quad \Omega_{\text{DM}} = \Omega_{\nu_R} h^2 + \Omega_{\bar{\nu}_R} h^2 = 0.12$

All data can be satisfied

Mass Spectrum



Mass Spectrum



Z_2 -even sector is like NMSSM!

Mass splitting between $H^\pm$ and H/A is caused by the mixing between H/A and N , which is necessary for DM annihilation
We can distinguish from **MSSM**

Effect of light Z_2 -odd particles deviate coupling of the Higgs h

$hVV, h\gamma\gamma, hhh, \dots$

We can distinguish from **NMSSM**

Mass Spectrum

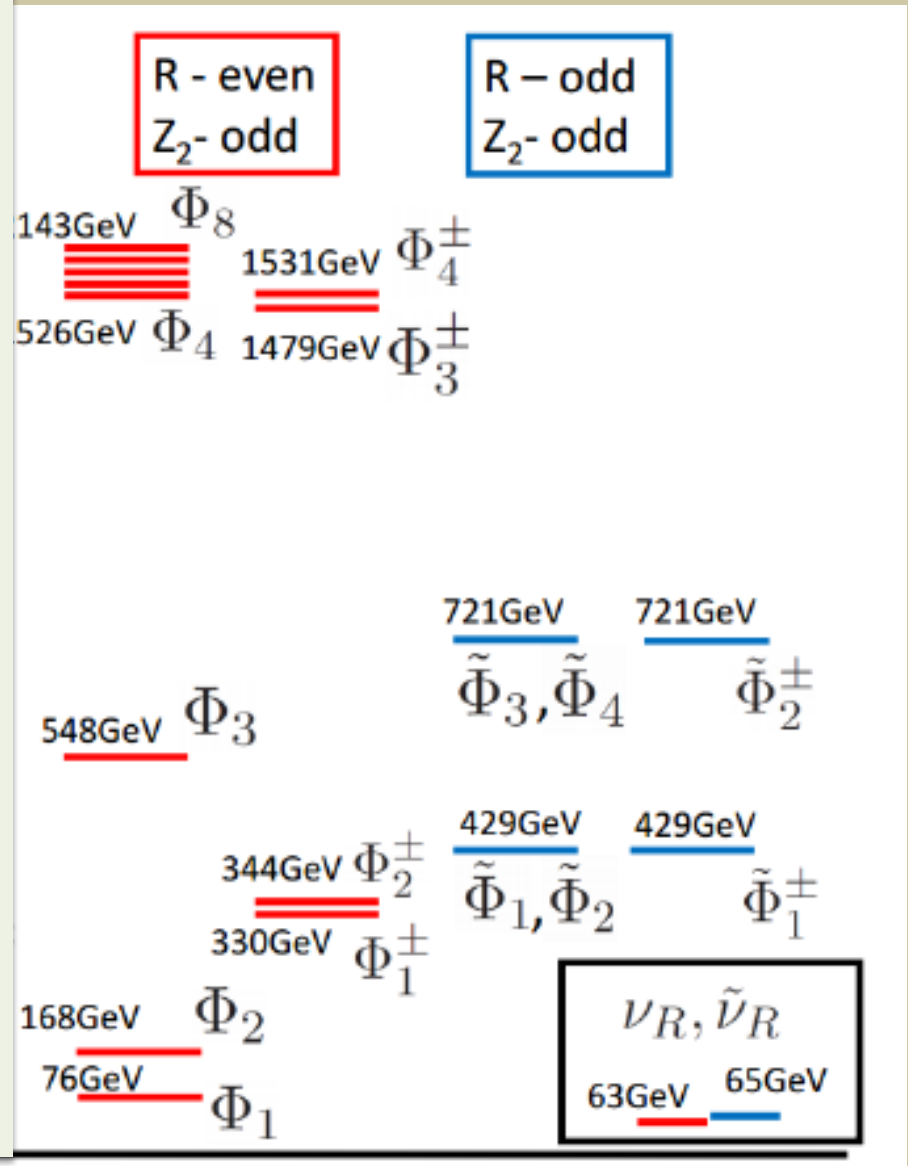
Z_2 -odd sector is rich

Light neutral/charged Z_2 -odd
Particle is required to realize
1st order EW phase transition

Multi component **DM** from
this sector

$$\nu_R, \tilde{\nu}_R$$

Z_2 -odd scalars are all used to
make 1 or 3 loop diagrams
for **neutrino masses**
(Ma and AKS type)



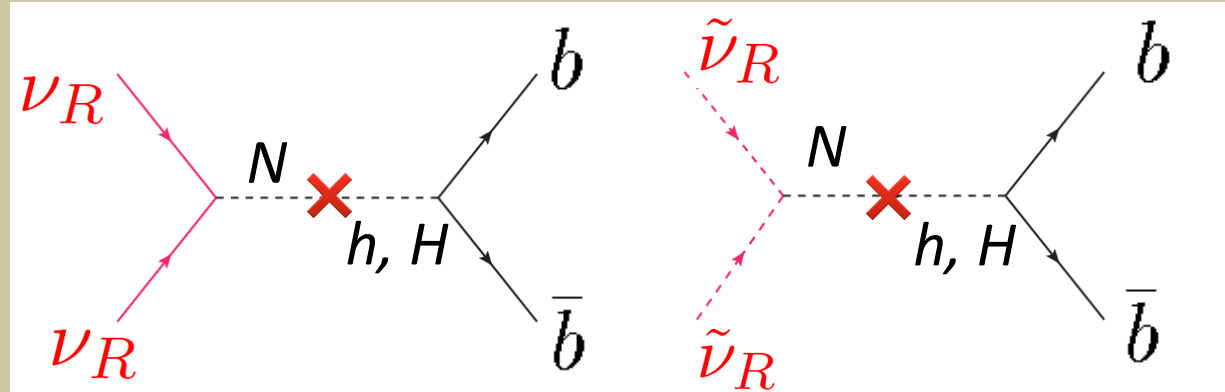
Dark Matter (RHN and RHSN)

Two discrete symmetries Z_2 and R :

$$(Z_2, R) = (-, +), (-, -)$$

$$\nu_R, \tilde{\nu}_R$$

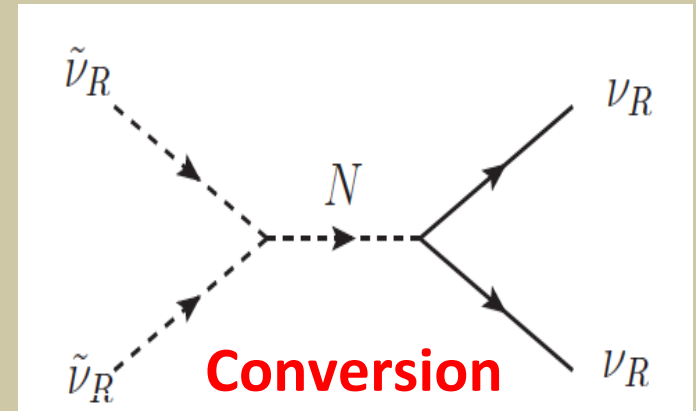
Multi-component DM



In our benchmark scenario,
the neutralino is not dark matter

$$\tilde{\chi}_1^0 \rightarrow \nu_R \tilde{\nu}_R$$

$$(+, -) \rightarrow (-, +) \oplus (-, -)$$



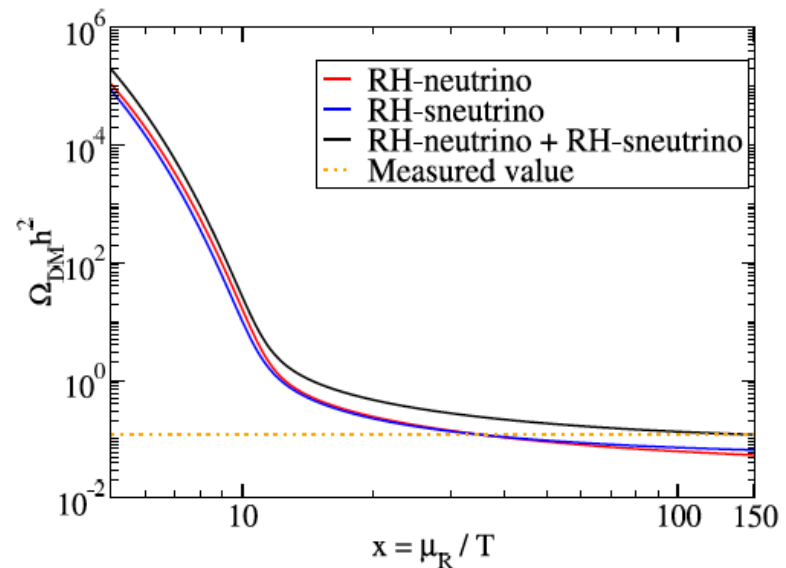
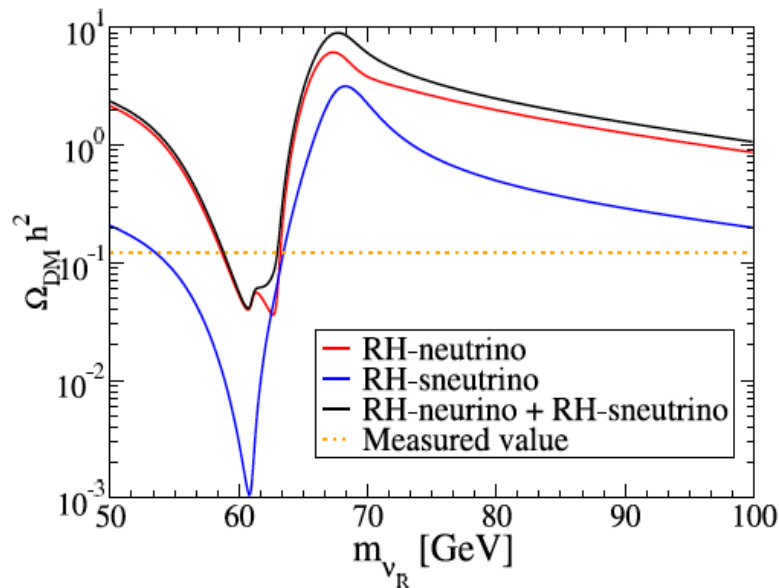
Relic Abundances of *RHN* and *RHSN*

Coupled Boltzmann Equations

$$\mu_R^{-1} = m_{\nu_R}^{-1} + m_{\tilde{\nu}_R}^{-1}$$

$$\frac{dY}{dx} = 0.264 g_*^{1/2} \left(\frac{\mu_R M_P}{x^2} \right) \left\{ -\langle \sigma_{\nu} v \rangle (Y^2 - Y_{\text{eq}}^2) - \langle \sigma_{\nu \tilde{\nu}} v \rangle \left(Y^2 - \tilde{Y}_{\text{eq}}^2 \frac{Y^2}{Y_{\text{eq}}^2} \right) + \langle \sigma_{\tilde{\nu} \nu} v \rangle \left(\tilde{Y}^2 - Y^2 \frac{\tilde{Y}_{\text{eq}}^2}{Y_{\text{eq}}^2} \right) \right\}$$

$$\frac{d\tilde{Y}}{dx} = 0.264 g_*^{1/2} \left(\frac{\mu_R M_P}{x^2} \right) \left\{ -\langle \sigma_{\tilde{\nu}} v \rangle (\tilde{Y}^2 - \tilde{Y}_{\text{eq}}^2) - \langle \sigma_{\nu \tilde{\nu}} v \rangle \left(\tilde{Y}^2 - Y^2 \frac{\tilde{Y}_{\text{eq}}^2}{Y_{\text{eq}}^2} \right) + \langle \sigma_{\nu \tilde{\nu}} v \rangle \left(\tilde{Y}^2 - Y^2 \frac{\tilde{Y}_{\text{eq}}^2}{Y_{\text{eq}}^2} \right) \right\}$$



For details of DM physics, see Naoki Machida's Talk

Summary of the Model



SUSY $SU(2)_H$ Fund. Rep. with 3 flavors

Right-Handed Neutrino

Simple

$T_1, T_2, T_3, T_4, T_5, T_6$

N_R

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**Confinement
at $O(10)$ TeV**

$$H = \langle T_i T_j \rangle$$

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Higgses = Mesons

IR

Extended
SUSY Higgs
sector with ν_R

$H_u, H_d, \Phi_u, \Phi_d, \dots$

Z_2 -odd

$\nu_R, \tilde{\nu}_R$



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$\nu_R, \tilde{\nu}_R$

Z_2 -odd

Rich!

Electroweak
Baryogenesis

Radiative
Neutrino Mass

WIMP
Dark Matter

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SUSY Higgs
sector with ν_R

$H_u, H_d, \Phi_u, \Phi_d, \dots$

$\nu_R, \tilde{\nu}_R$

Z_2 -odd

Rich!

Electroweak
Baryogenesis

Radiative
Neutrino Mass

WIMP
Dark Matter

All additional scalars contribute to solve the problems

How test the model at colliders

- **Direct search** of exotic charged scalars and fermions at LHC (expected for 14 TeV Run) and ILC

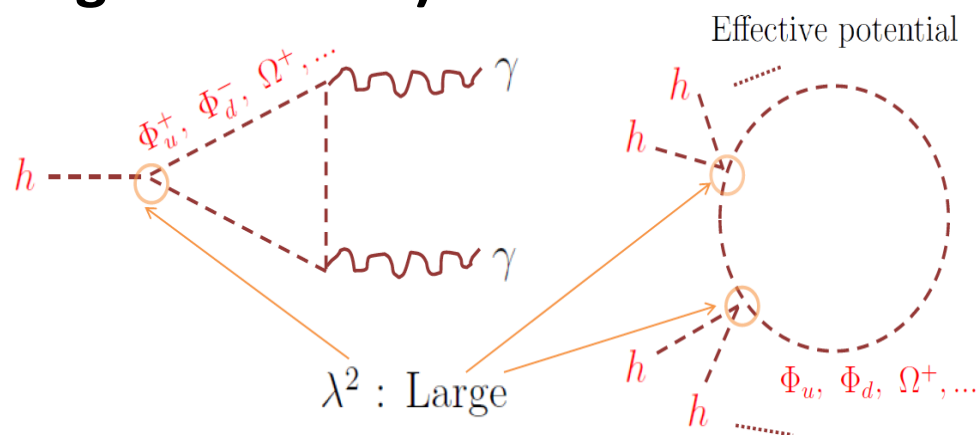
- **Indirect Signatures on Higgs coupling**
(essence of Strong-but-Light scenario)

- Non-decoupling effect
- Deeply connected with $\varphi c/Tc > 1$

**Large deviations
in the $h\gamma\gamma$, hhh coupling!**

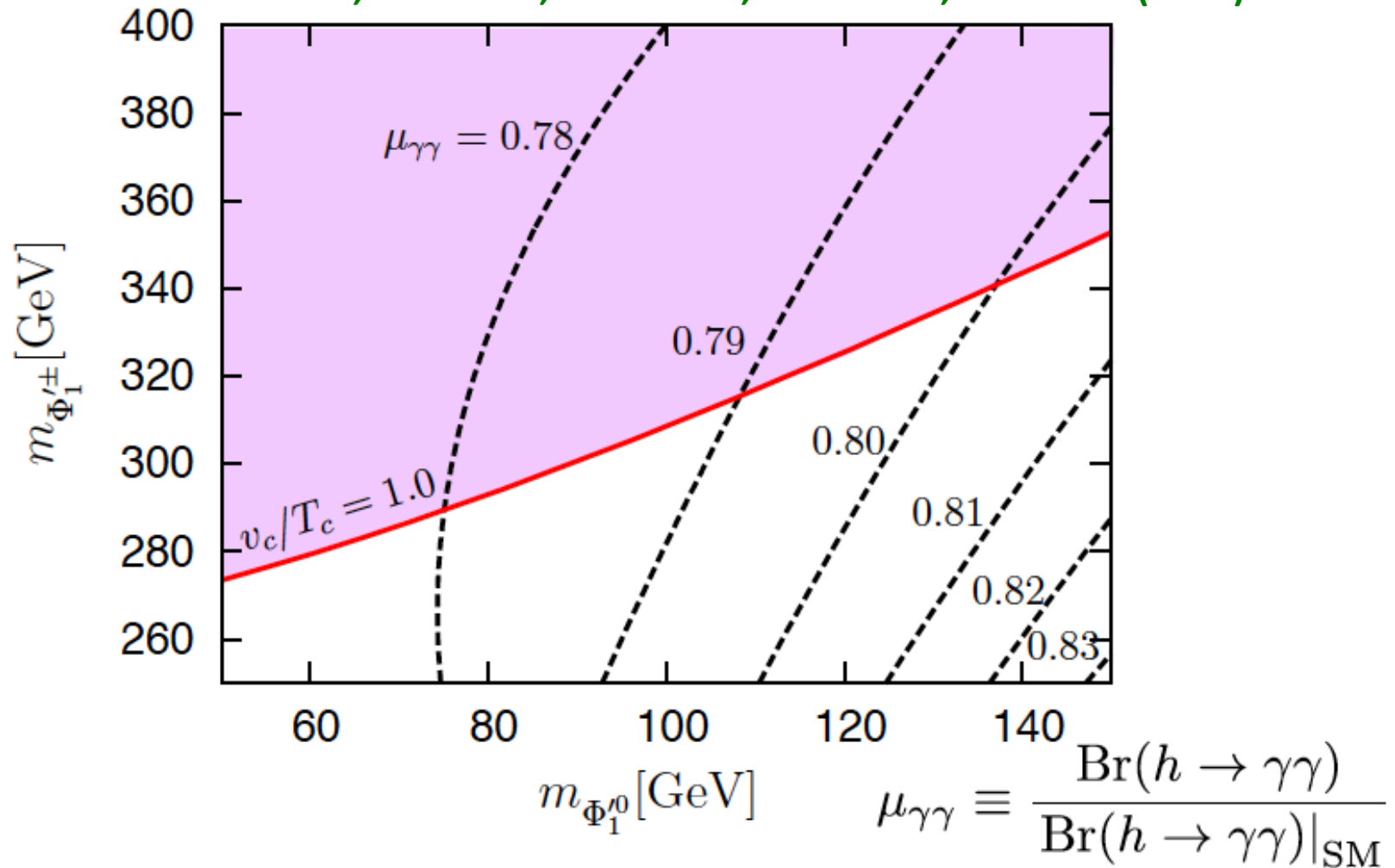
Coupling measurement

- hWW , hff $O(1)\%$ at ILC
- $h\gamma\gamma$ coupling $O(10)\%$ at HL-LHC and ILC500
- hhh coupling $O(10)\%$ at ILC 500-1000



Contribution to $h\gamma\gamma$

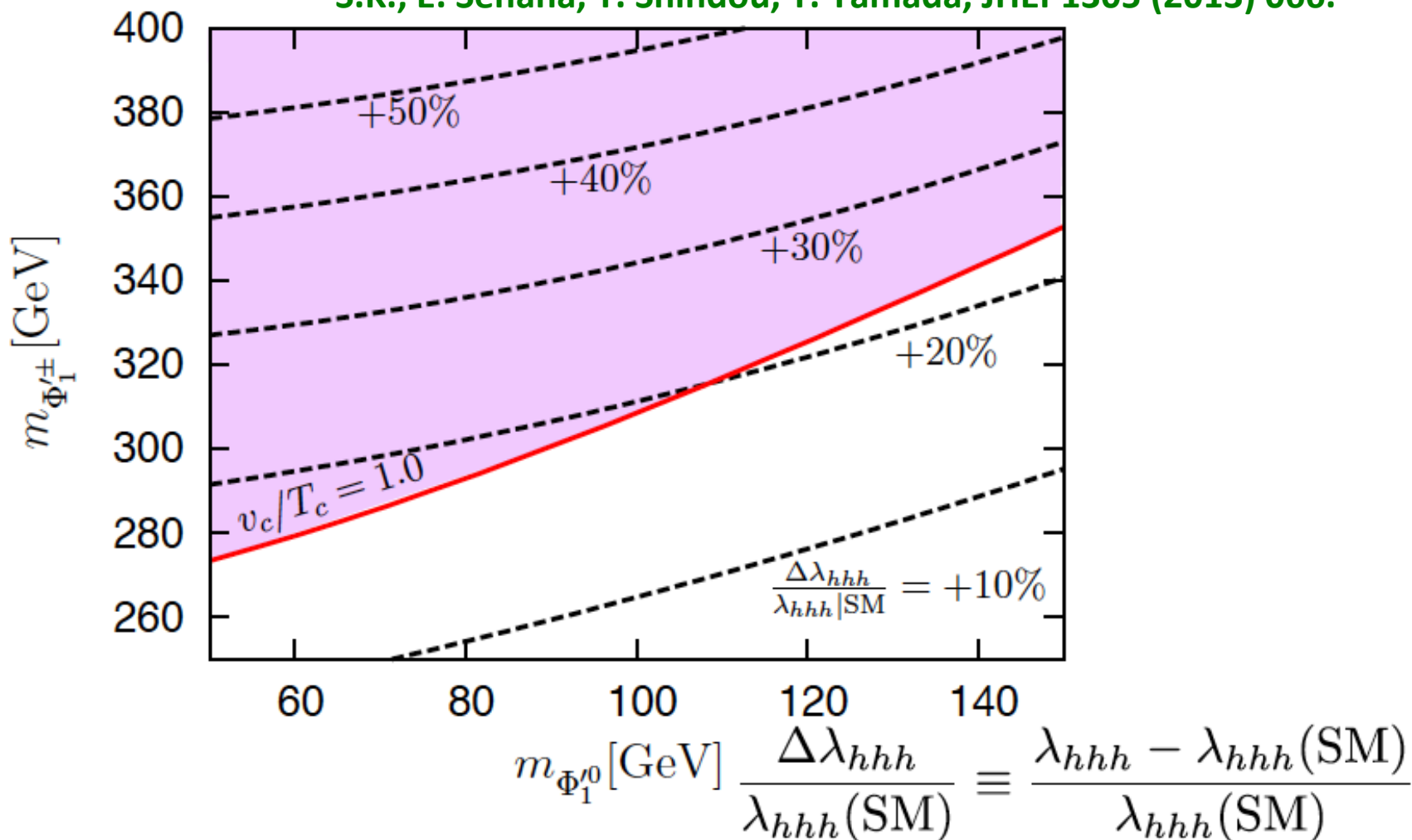
S.K., E. Senaha, T. Shindou, T. Yamada, JHEP1305 (2013) 066.



< 20 % deviation is possible in the region of $v_c/T_c > 1$

hhh coupling

S.K., E. Senaha, T. Shindou, T. Yamada, JHEP1305 (2013) 066.



> 20 % deviation is possible in the region of $v_c/T_c > 1$

Conclusion

- We discussed SUSY $SU(2)_H$ gauge theory with 3 flavour with Z_2 parity
- Confinement occurs at Λ_H , below which the effective theory contains the Higgs sector with 4 Doublets + many singlets with only one λ coupling
Higgses as Mesons [Particle content (15 mesons) for radiative seesaw]
- Imposing the condition of EW 1st OPT on λ , it turns out that Λ_H is $O(10)$ TeV
- Introducing a Z_2 -odd RH-Neutrino, 1- and 3-loop radiative seesaw scenarios can be simultaneously realized at the TeV scale, where all new scalars are used.
(Hybrid radiative seesaw scenario)
- In our model, Strong 1st OPT, Neutrino masses and DM are explained simultaneously.
A benchmark scenario where all relevant data are satisfied.
- Our model can be tested at colliders via many discriminative properties (large deviations in $h\gamma\gamma$ and $h\gamma h$ couplings, direct searches of exotic scalar fields, multi-component DM, LFV, ...)

Summary of the Model

UV

SUSY $SU(2)_H$ Fund. Rep. with 3 flavors

Right-Handed Neutrino

Simple

$T_1, T_2, T_3, T_4, T_5, T_6$

N_R

Z_2 -odd assigned

Confinement
at $O(10)$ TeV

$$H = \langle T_i T_j \rangle$$

Higgses = Mesons

IR

Extended
SUSY Higgs
sector with ν_R

$H_u, H_d, \Phi_u, \Phi_d, \dots$

$\nu_R, \tilde{\nu}_R$

Z_2 -odd

Rich!

Electroweak
Baryogenesis

Radiative
Neutrino Mass

WIMP
Dark Matter

All additional scalars contribute to solve the problems

Advertisement!

We organize
the 2nd Workshop on
“Higgs as a Probe of New
Physics 2015 (HPNP2015)”
at Univ. of Toyama, Japan
11-15 February 2015

All of you are
cordially invited !

HPNP2015

The 2nd Toyama International Workshop on
“Higgs as a Probe of New Physics 2015”

February 11-15, 2015, University of Toyama, Japan

Invited Speakers

Abdesslam Arhrib* (Abdelmalek Essaadi U.)	Edmond L. Berger (ANL)
Fawzi Boudjema* (Annecy, LAPTH)	Chengwei Chiang (NCU)
Eung-Jin Chun (KIAS)	Abdelhak Djouadi (LPT, Orsay)
Keisuke Fujii (KEK/SOKENDAI)	Christophe Grojean (U. of Barcelona)
Howard E. Haber* (UC Santa Cruz)	Kaoru Hagiwara (KEK/SOKENDAI)
Sven Heinemeyer (IFCA)	Pyungwon Ko (KIAS)
Maria Krawczyk (U. of Warsaw)	Ernest Ma (UC. Riverside)
Stefano Moretti (U. of Southampton)	Salah Nasri (UAE U. & Oran U.)
Michael Peskin (SLAC)	Rui Santos (ISEL, U. of Lisbon)
Reisaburo Tanaka (LAL, Orsay)	C.-P. Yuan (Michigan State U.)

Public Lecture (15th February, Afternoon)

Hitoshi Murayama (Kavli IPMU / UC Berkeley)

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Chengwei Chiang (NCU)	Abdelhak Djouadi (LPT, Orsay)
Howard E. Haber (UC Santa Cruz)	Yutaka Hosotani (Osaka U.)
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URL : <http://hodo.sci.u-toyama.ac.jp/theory/HPNP2015/>

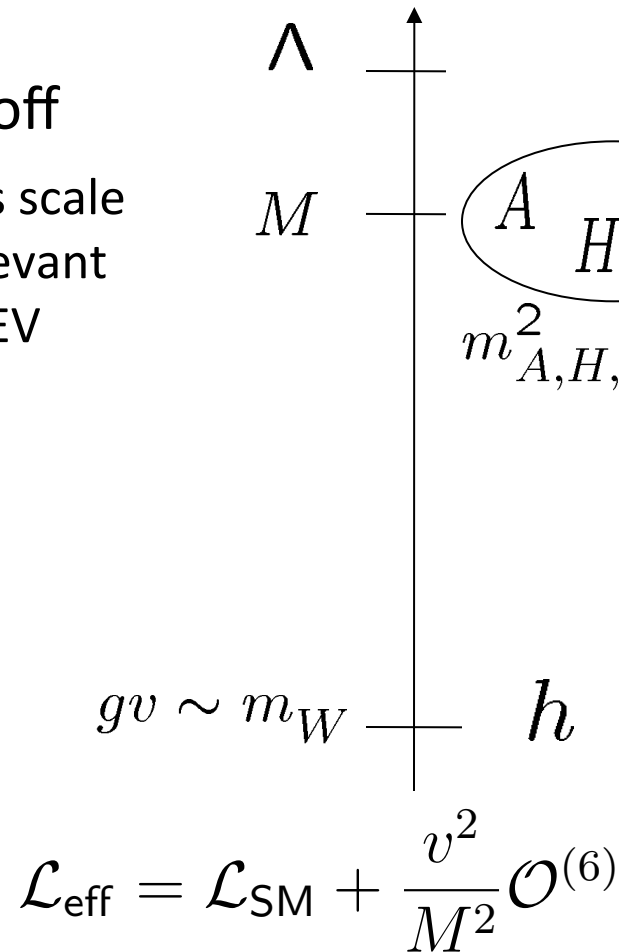


Back Up Slides

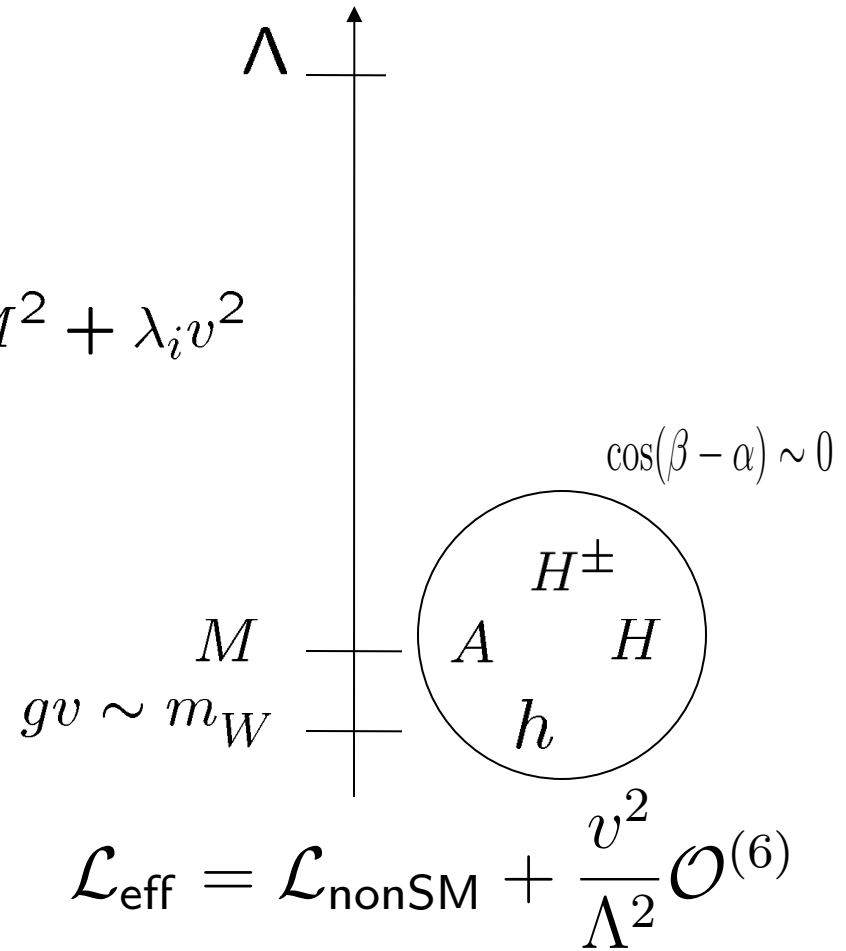
Two Possibilities

Λ : Cutoff

M : Mass scale
irrelevant
to VEV



Effective Theory is the SM



Effective Theory is an extended Higgs sector

Non-decoupling effect

EW Baryogenesis and the hhh coupling

Higgs Potential at Finite Temperatures

$$V_T(\phi, T) = D(T^2 - T_0^2)\phi^2 - ET\phi^3 + \frac{\lambda_T}{4}\phi^4 + \dots$$

$$\phi_c/T_c = 2E/\lambda_{T_c}$$

$$E = \frac{1}{12\pi v^3}(6m_W^3 + 3m_Z^3) + \text{Non-decoupling effect of new particles}$$

$$\lambda_T = m_h^2/2v^2 + \log \text{ corrections}$$

Condition of strongly 1st OPT

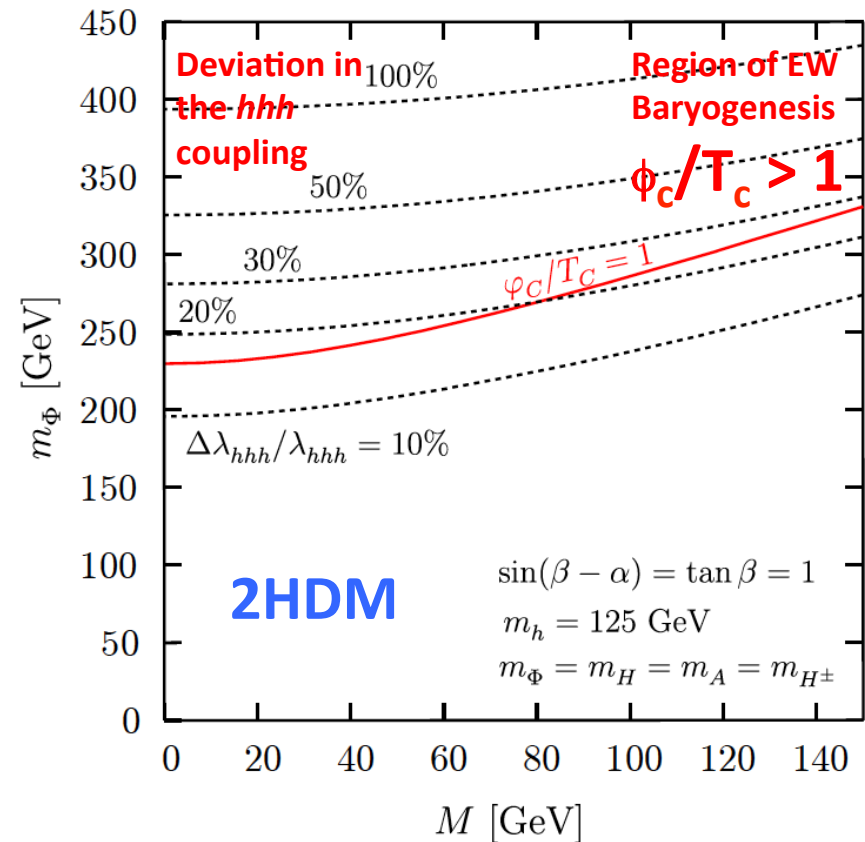
$$\phi_c/T_c > 1 \Rightarrow 2E/\lambda_{T_c} > 1$$

Strongly 1st OPT

$\Leftrightarrow$ Non-decoupling effect

$\Leftrightarrow$ large deviation in hhh

SK, Okada, Senaha (2005)



If hhh can be measured by 10 % at ILC, EW Baryogenesis can be tested

Connection between cosmology and collider physics

2 Higgs Doublet Model

$$V_{\text{THDM}} = +m_1^2 |\Phi_1|^2 + m_2^2 |\Phi_2|^2 - \underline{m_3^2 (\Phi_1^\dagger \Phi_2 + \Phi_2^\dagger \Phi_1)} \\ + \frac{\lambda_1}{2} |\Phi_1|^4 + \frac{\lambda_2}{2} |\Phi_2|^4 + \lambda_3 |\Phi_1|^2 |\Phi_2|^2 \\ + \lambda_4 |\Phi_1^\dagger \Phi_2|^2 + \frac{\lambda_5}{2} \left[(\Phi_1^\dagger \Phi_2)^2 + (\text{h.c.}) \right]$$

$$\Phi_1 \text{ and } \Phi_2 \Rightarrow \underbrace{h, H, A^0, H^\pm}_{\substack{\uparrow \text{CPEven} \quad \uparrow \text{CPodd} \\ \uparrow \text{charged}}} \oplus \text{Goldstone bosons}$$

$$m_h^2 = v^2 \left(\lambda_1 \cos^4 \beta + \lambda_2 \sin^4 \beta + \frac{\lambda}{2} \sin^2 2\beta \right) + \mathcal{O}\left(\frac{v^2}{M_{\text{soft}}^2}\right),$$

$$m_H^2 = M_{\text{soft}}^2 + v^2 (\lambda_1 + \lambda_2 - 2\lambda) \sin^2 \beta \cos^2 \beta + \mathcal{O}\left(\frac{v^2}{M_{\text{soft}}^2}\right),$$

$$m_{H^\pm}^2 = M_{\text{soft}}^2 - \frac{\lambda_4 + \lambda_5}{2} v^2,$$

$$m_A^2 = M_{\text{soft}}^2 - \lambda_5 v^2.$$

M_{soft} : soft breaking scale

$$\Phi_i = \begin{bmatrix} w_i^+ \\ \frac{1}{\sqrt{2}}(h_i + v_i + i a_i) \end{bmatrix} \quad (i = 1, 2)$$

Diagonalization

$$\begin{bmatrix} h_1 \\ h_2 \end{bmatrix} = \begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix} \begin{bmatrix} H \\ h \end{bmatrix} \quad \begin{bmatrix} z_1^0 \\ z_2^0 \end{bmatrix} = \begin{bmatrix} \cos \beta & -\sin \beta \\ \sin \beta & \cos \beta \end{bmatrix} \begin{bmatrix} z^0 \\ A^0 \end{bmatrix} \\ \begin{bmatrix} w_1^\pm \\ w_2^\pm \end{bmatrix} = \begin{bmatrix} \cos \beta & -\sin \beta \\ \sin \beta & \cos \beta \end{bmatrix} \begin{bmatrix} w^\pm \\ H^\pm \end{bmatrix}$$

$$\frac{v_2}{v_1} \equiv \tan \beta$$

$$M_{\text{soft}} \left(= \frac{m_3}{\sqrt{\cos \beta \sin \beta}} \right):$$

soft-breaking scale
of the discrete symm.

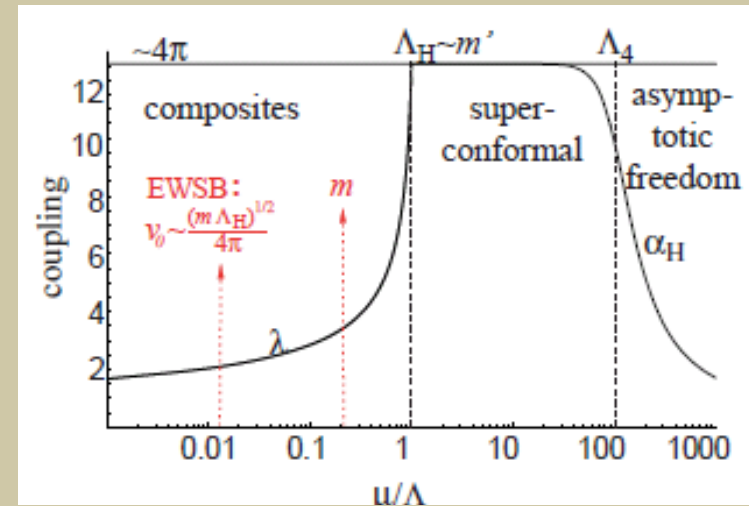
What is the UV Theory beyond Λ_{cutoff} ?

Hint) Minimal SUSY Fat Higgs Model

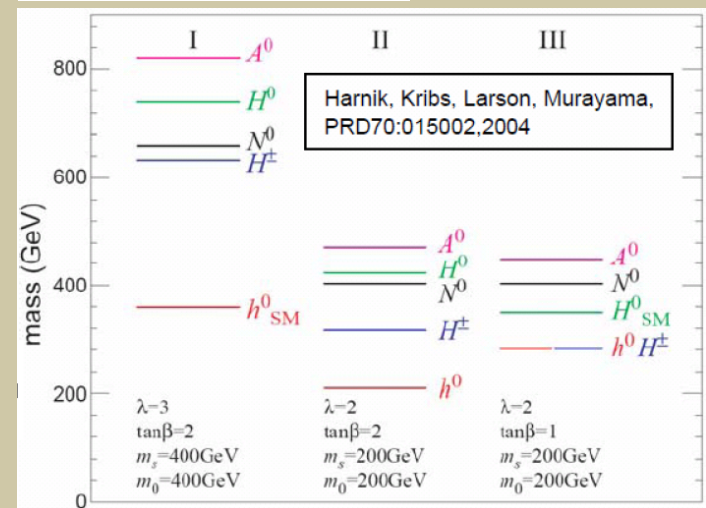
Harnik, Kribs, Larson, Murayama, 2004

- $SU(2)_H$ gauge theory with $N_f = 4 \rightarrow 3$
- Confinement at the cutoff Λ_H ($N_f=3$)
- Below Λ_H , Higgs fields appear as composite states
- Low energy effective theory is minimized to be the **NMSSM**
(2 doublets) H_1, H_2 , (1 singlet) N
 - In the $SU(2)_H$ model with $N_f=3$, **15 composite superfields** appear
 - Unnecessary 10 composite fields are made heavy **by introducing additional heavy fields**
- SM-like Higgs boson is heavy (fat)

$$m_h^2 \simeq \lambda^2 v^2 + \mathcal{O}(m_Z^2)$$



$$W = \lambda(NH_1H_2 - v_0^2)$$



Like the minimal SUSY Fat Higgs

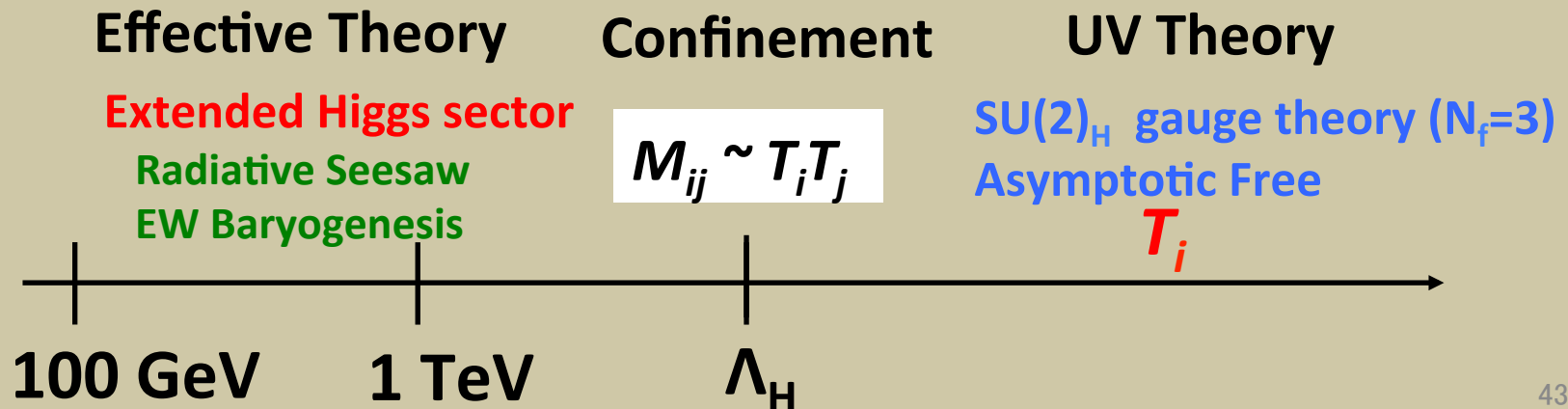
We here consider the model based on the SUSY $SU(2)_H$ gauge theory with $N_f=3$ in order **to explain BSM phenomena**

Neutrino masses

Dark matter

Baryon Asymmetry of the Universe

by the TeV scale physics (below Λ_H)



Deriving Yukawa Couplings

- SM Yukawa couplings can be generated by introducing elementary $SU(2)_L$ doublets, H'_u, H'_d , that couple as

$$W_{Yuk} = (T_1, T_2)T_3 H'_d + (T_1, T_2)T_4 H'_u + \underbrace{m H'_u H'_d}_{m \sim \Lambda_H} + y_{u ij} Q^i H'_u U^j + y_{d kl} Q^k H'_d D^l + \dots$$

and integrating them out below the scale $m (\sim \Lambda_H)$.

- Higgs = composite superfield

Top quark = elementary superfield

 Difficulty in deriving $O(1)$ top Yukawa coupling

- But we already have an elegant mechanism for this.

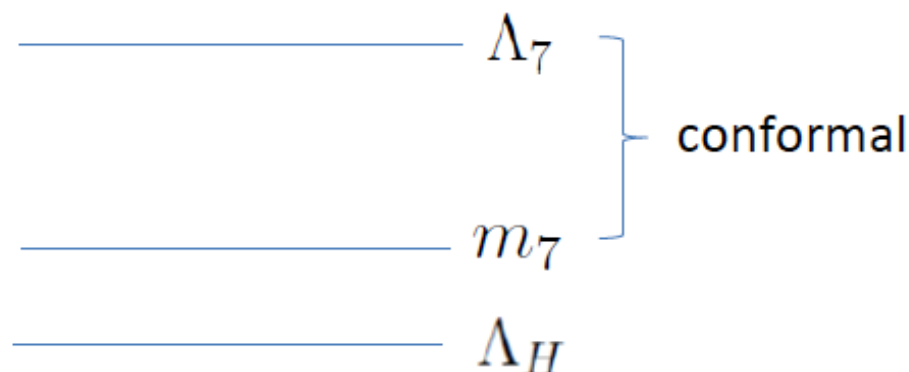
Conformal Enhancement

Murayama

Introduce two more $SU(2)_H$ doublets, T_7, T_8 , with mass term: $W_7 = m_7 T_7 T_8$ ($m_7 > \Lambda_H$).

The theory above the scale m_7 is in the conformal window.

Assume that the theory approaches to the IR fixed point at the scale $\Lambda_7 (> m_7)$.



➡ Yukawa couplings are enhanced by $\left(\frac{\Lambda_7}{m_7}\right)^{1/2}$ while running from Λ_7 to m_7 .

Snowmass White Paper (Aug. 2013)

Facility	LHC	HL-LHC	ILC500	ILC500-up	ILC1000	ILC1000-up	CLIC	TLEP (4 IPs)
$\sqrt{s}$ (GeV)	14,000	14,000	250/500	250/500	250/500/1000	250/500/1000	350/1400/3000	240/350
$\int \mathcal{L} dt$ (fb $^{-1}$)	300/expt	3000/expt	250+500	1150+1600	250+500+1000	1150+1600+2500	500+1500+2000	10,000+2600
κ_γ	5 – 7%	2 – 5%	8.3%	4.4%	3.8%	2.3%	–/5.5/<5.5%	1.45%
κ_g	6 – 8%	3 – 5%	2.0%	1.1%	1.1%	0.67%	3.6/0.79/0.56%	0.79%
κ_W	4 – 6%	2 – 5%	0.39%	0.21%	0.21%	0.13%	1.5/0.15/0.11%	0.10%
κ_Z	4 – 6%	2 – 4%	0.49%	0.24%	0.44%	0.22%	0.49/0.33/0.24%	0.05%
κ_ℓ	6 – 8%	2 – 5%	1.9%	0.98%	1.3%	0.72%	3.5/1.4/<1.3%	0.51%
κ_d	10 – 13%	4 – 7%	0.93%	0.51%	0.51%	0.31%	1.7/0.32/0.19%	0.39%
κ_u	14 – 15%	7 – 10%	2.5%	1.3%	1.3%	0.76%	3.1/1.0/0.7%	0.69%

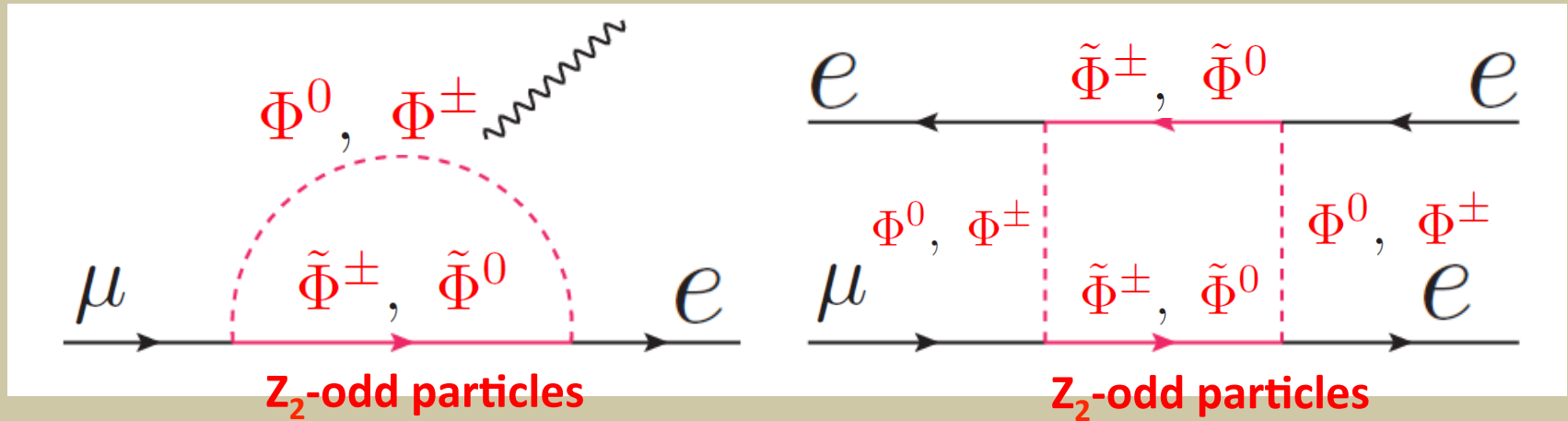
$$g(hxx)=\kappa_x g(hxx)_{SM}$$

ILC Higgs White Paper
arXiv:1310.0763

*Asner, Barklow, Fujii,
Haberman, Kanemura,
Miyamoto, Weiglein,
et al.*

	ILC(250)	ILC(500)	ILC(1000)	ILC(LumUp)
$\sqrt{s}$ (GeV)	250	250+500	250+500+1000	250+500+1000
L (fb $^{-1}$)	250	250+500	250+500+1000	1150+1600+2500
$\gamma\gamma$	17 %	8.3 %	3.8 %	2.3 %
gg	6.1 %	2.0 %	1.1 %	0.7 %
WW	4.7 %	0.4 %	0.3 %	0.2 %
ZZ	0.7 %	0.5 %	0.5 %	0.3 %
$t\bar{t}$	6.4 %	2.5 %	1.3 %	0.9 %
$b\bar{b}$	4.7 %	1.0 %	0.6 %	0.4 %
$\tau^+\tau^-$	5.2 %	1.9 %	1.3 %	0.7 %
$\Gamma_T(h)$	9.0 %	1.7 %	1.1 %	0.8 %
$\mu^+\mu^-$	91 %	91 %	16 %	10 %
hhh	–	83 %	21 %	13 %
BR(invis.)	< 0.7 %	< 0.7 %	< 0.7 %	< 0.3 %
$c\bar{c}$	6.8 %	2.9 %	2.0 %	1.1 %

Lepton Flavor Violation



$$\text{Br}(\mu \rightarrow e \gamma) = 2.8 \times 10^{-13} \quad (< 5.7 \times 10^{-13})$$

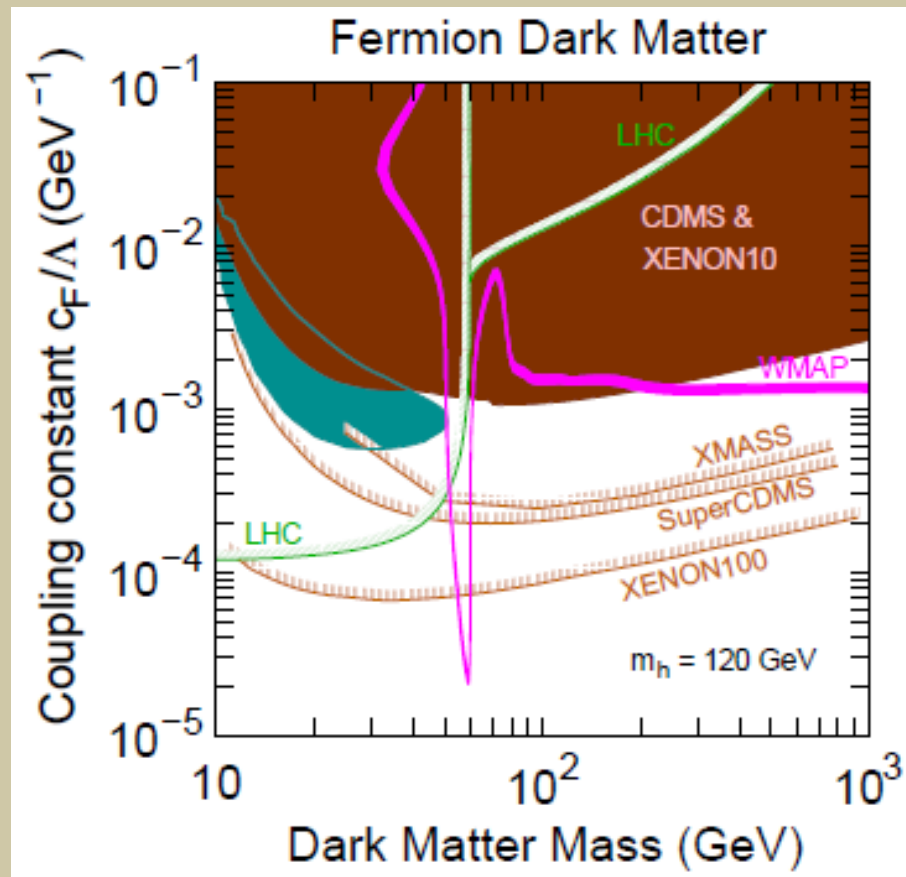
$$\text{Br}(\mu \rightarrow e e e) = 1.6 \times 10^{-16} \quad (< 1.0 \times 10^{-12})$$

Current data

MEG: arXiv:1303.0754v2 [hep-ex]

PDG, Phys. Rev. D86, 010001 (2012)

Case of the SM-Higgs Portal DM



SK, S. Matsumoto, T. Nabeshima, N. Okada, 2010

In 2012 July, a boson was found

The mass is 126 GeV

Spin/Parity 0^+

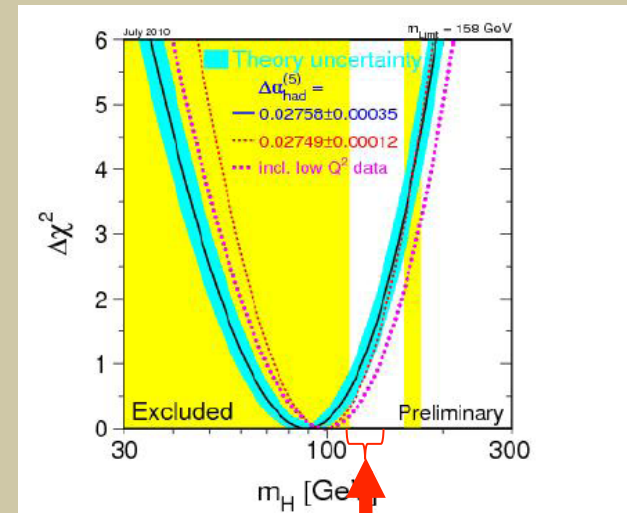
It couples to $\gamma\gamma$, WW , ZZ , bb , $\tau\tau$, ...

This is really a Higgs!

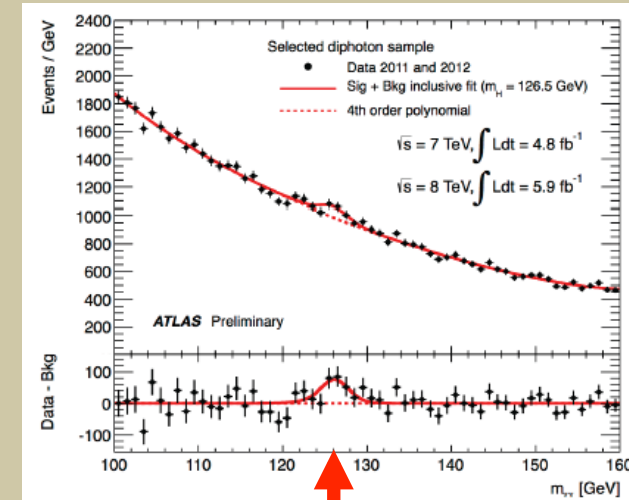


Measured Couplings look consistent with the SM Higgs

Why SM-like???



Higgs Mass indicated by LEP/SLC



ATLAS/CMS July 2012

New Particle !

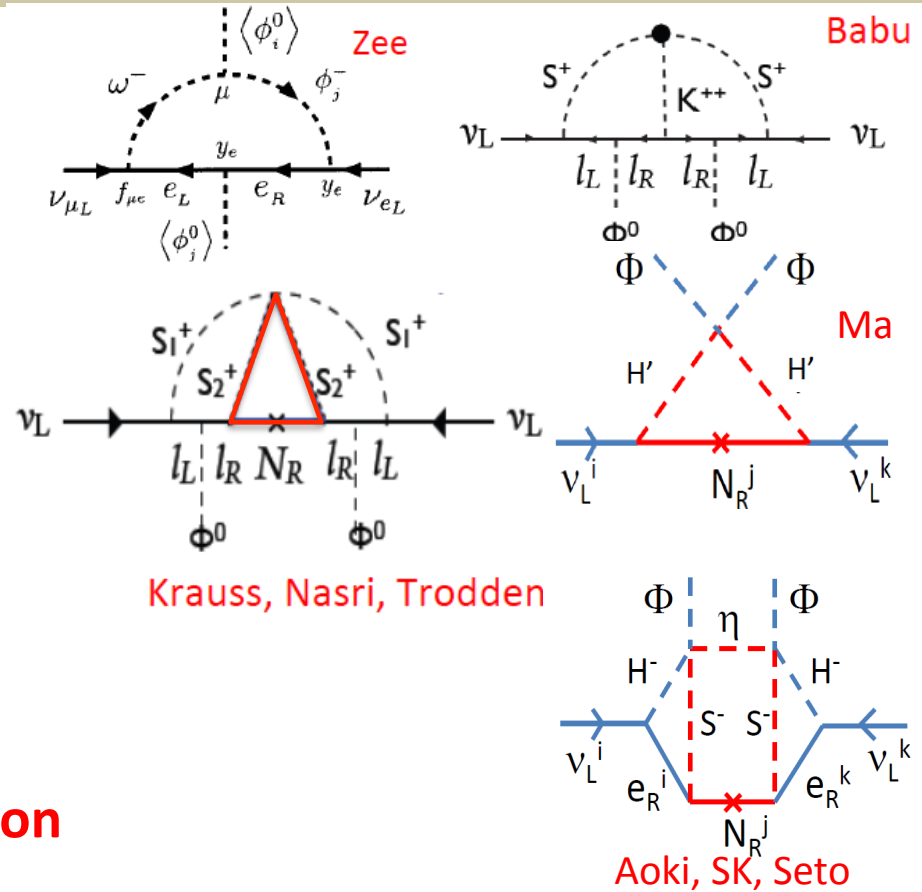
Explanations by the TeV scale physics

Radiative Seesaw Scenario

- Extended Higgs sector
- Z_2 parity
 - **Neutrino mass** generated at loop levels
 - **WIMP Dark Matter**
 - Lightest Z_2 -odd particle
 - LSP (in SUSY extension)

Electroweak Baryogenesis

- Sphaleron
- **Additional CP Phases**
- **Strong 1st Order Phase Transition**



These scenarios are strongly related to the Higgs physics!