

Multi-Phase Dynamical Symmetry Breaking by Scalar Dark Matter

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- Classical scale invariance can relieve the hierarchy problem
- Both dilaton and Higgs can be light near the critical phase boundary
- Quantum corrections driven by couplings to dark matter

3 Works

- Multi-phase dynamical symmetry breaking K.K., Luca Marzola, Martti Raidal, Alessandro Strumia
 Phys.Lett.B 816 (2021) 136241 [arXiv:2102.01084]
- Model-independent collider phenomenology Katri Huitu, K.K., Niko Koivunen, Luca Marzola, Subhadeep Mondal, Martti Raidal

Phys.Rev.D 105 (2022) 9, 095036 [arXiv:2201.00824]

 Scalar singlet dark matter
K.K., Niko Koivunen, Aleksei Kubarski, Luca Marzola, Martti Raidal, Alessandro Strumia & Venno Vipp

Phys.Lett.B 832 (2022) 137214 [arXiv:2204.01744]

4 Dynamical Symmetry Breaking

- No dimensionful terms in the scalar potential
- Classical scale invariance broken by quantum corrections
- Minimum near the flat direction (V = 0)

S. R. Coleman, E. J. Weinberg, Phys. Rev. D 7 (1973) 1888–1910.

E. Gildener, S. Weinberg, Phys. Rev. D 13 (1976) 3333.



5 Model Tree-level scalar potential given by

$$V = \frac{1}{4}\lambda_{H}h^{4} + \frac{1}{4}\lambda_{S}s^{4} + \frac{1}{4}\lambda_{HS}h^{2}s^{2} + \frac{1}{4}\lambda_{HS'}h^{2}s^{2} + \frac{1}{4}\lambda_{SS'}s^{2}s'^{2} + \frac{1}{4}\lambda_{S'}s'^{4}$$

- In addition to the Higgs boson h, dilaton s, and dark matter s'
- Invariant under a $\mathbb{Z}_2 \otimes \mathbb{Z}'_2$ with $s \to -s$ and $s' \to -s'$
- VEVs $\langle h \rangle \equiv v$ and $\langle s \rangle \equiv w$
- Sufficient to require λ_{s'}, λ_{Hs'}, λ_{ss'} > 0 to avoid a VEV for dark matter s'

6 Phases

Symmetry is broken dynamically when a combination of quartics runs through zero

h)
$$\lambda_H = 0$$
 and $\lambda_S, \lambda_{HS} > 0$
for $v \neq 0$ and $w = 0$
s) $\lambda_S = 0$ and $\lambda_H, \lambda_{HS} > 0$
for $w \neq 0$ and $v = 0$
sh) $\lambda_{HS} = -2\sqrt{\lambda_H\lambda_S} < 0$
and $\lambda_H, \lambda_S > 0$
for $w, v \neq 0$



7 Phases

The sh phase gives, in the Gildener-Weinberg approximation,

$$\frac{v}{w} = \sqrt{\frac{-\lambda_{HS}}{2\lambda_{H}}}$$

- Gildener-Weinberg approximation breaks down in the critical regime where the s and sh phases coincide
- Better approximation takes into account running between the flat direction and the minimum

- In the Gildener-Weinberg regime, $\lambda_{HS}(w) \approx \lambda_{HS}(s_{flat})$
- In the multi-phase critical regime, $\lambda_{HS}(w) \ll \lambda_{HS}(s_{flat})$

9 Gildener-Weinberg regime

Model studied in the Gildener-Weinberg regime in:

Ishiwata, Phys. Lett. B 710 (2012) 134–138 [arXiv:1112.2696] Gabrielli, Heikinheimo, K.K, Racioppi, Raidal, Spethmann, Phys. Rev. D 89 (1) (2014) 015017 [arXiv: 1309.6632]

Kang, Zhu, Phys. Rev. D 102 (5) (2020) 053011 [arXiv:2003.02465]

10 Multi-Phase Critical Regime

In the multi-phase critical limit, the two phases, s and *sh*, merge:

$$\lambda_{\rm S}(\bar{\mu})=0,\qquad\lambda_{\rm HS}(\bar{\mu})pprox 0$$

- Higgs massless at $\lambda_{HS}(\bar{\mu}) = 0$
- Another field to drive the running

Quantum Corrections

In the critical limit, the dominant mass is

$$m_{s'}^2 = \frac{1}{2}(\lambda_{SS'}s^2 + \lambda_{HS'}h^2)$$

• β -functions given by

$$\beta_{\lambda_{HS}} \approx \frac{1}{2} \lambda_{SS'} \lambda_{HS'}, \qquad \beta_{\lambda_S} \approx \frac{1}{4} \lambda_{SS'}^2$$

 Dark matter Higgs portal λ_{HS'} important for symmetry breaking and direct detection

12 Quantum Corrections

Effective potential

$$V_{\rm eff} = V^{(0)} + V^{(1)} \approx V^{(0)}(\lambda_i(\mu = s))$$

Couplings run as

$$\lambda_{S}^{\text{eff}}(s) = \frac{\beta_{\lambda_{S}}}{(4\pi)^{2}} \ln \frac{s^{2}}{s_{S}^{2}},$$
$$\lambda_{HS}^{\text{eff}}(s) = \frac{\beta_{\lambda_{HS}}}{(4\pi)^{2}} \ln \frac{R s^{2}}{e^{-1/2} s_{S}^{2}}$$

13 Quantum Corrections

■ Flat direction scale s_{flat} ≈ s_S
■ Dilaton VEV w ≈ s_S e^{-1/4}
■ Higgs VEV

$$v = \frac{w}{4\pi} \sqrt{-\frac{\beta_{\lambda_{HS}} \ln R}{2\lambda_{H}}}$$

suppressed by quantum corrections





15 Parameters

- Higgs mass $m_h \approx \sqrt{2\lambda_H} v \approx 125.1 \text{ GeV}$
- Higgs VEV v = 246.2 GeV
- Free parameters: m_{s} , $m_{s'}$, $\ln R$
- $\lambda_{S'}$ largely irrelevant (not too large)

16 Parameters

$$\begin{split} m_h^2 &\approx -\frac{\beta_{\lambda_{HS}}}{(4\pi)^2} w^2 \ln R \approx 2\lambda_H v^2, \\ m_s^2 &\approx 2 \frac{\beta_{\lambda_s}}{(4\pi)^2} w^2, \\ m_{s'}^2 &\approx \frac{1}{2} \lambda_{SS'} w^2, \end{split}$$

while the mixing angle is given by

$$\theta \approx \frac{m_{hs}^2}{m_s^2 - m_h^2} = \frac{\beta_{\lambda_{Hs}}(1 + \ln R)}{2\beta_{\lambda_s} + \beta_{\lambda_{Hs}} \ln R} \frac{v}{w},$$

where m_{hs}^2 is the mixing mass term among h and s

17 Parameters

$$\begin{split} \lambda_{SS'} &\approx \frac{(4\pi)^2 m_s^2}{m_{s'}^2}, \\ \lambda_{HS'} &\approx -\frac{(4\pi)^2 m_h^2}{m_{s'}^2 \ln R}, \\ w &\approx \frac{\sqrt{2}m_{s'}^2}{4\pi m_s} \end{split}$$

and

$$\theta \approx \frac{2\sqrt{2}\pi m_s m_h^2 v(1 + \ln R)}{(m_h^2 - m_s^2) m_{s'}^2 \ln R}$$

18 Dark Matter Abundance

 $\Omega_{\rm DM}h^2 = 0.120 \pm 0.001$

N. Aghanim, et al., Astron. Astrophys. 641 (2020) A6

■ Heavy-DM limit m_{s'} ≫ m_s, m_h, where the non-relativistic DM annihilations cross section is simply given by

$$\sigma_{\rm ann} v_{\rm rel} \approx \frac{\lambda_{\rm SS'}^2 + 4\lambda_{\rm HS'}^2}{64\pi m_{s'}^2} \approx 4\pi^3 \frac{m_s^4 + 4m_h^4/\ln^2 R}{m_{s'}^6}.$$

Higgs-dominated if $m_s \ll m_h$, dilaton-dominated if $m_s \gg m_h$

19 Direct Detection

Effective coupling to nucleons

$$\frac{f_N m_N}{v} h \overline{N} N,$$

gives

$$\sigma_{\rm SI} \approx \frac{64\pi^3 f_N^2 m_N^4}{m_{s'}^6},$$

where $m_N = 0.946$ GeV is the nucleon mass and $f_N \approx 0.3$ 20 Direct Detection



21 Parameter Space: Bounds

- Perturbativity: $\lambda_{HS}, \lambda_{SS'} < 4\pi$
- Fit of Higgs couplings
- $h \rightarrow ss decay$
- Direct detection









26 Perturbativity



27 In Progress: Freeze-in, and More

- Type-I seesaw mechanism with right-handed neutrinos N_R
- Yukawa Lagrangian

$$-L_{Y} = y_{H} \bar{\ell} \tilde{H} N_{R} + \frac{y_{S}}{2} S \bar{N}_{R}^{c} N_{R} + h.c.,$$

where $\tilde{H} \equiv i\tau_2 H^*$

Inflation with the dilaton s

28 In Progress: Freeze-in and More



29 Conclusions

- Classically scale-invariant model with dynamical symmetry breaking driven by couplings to dark matter
- Crucial corrections to Gildener-Weinberg approximation in the multi-phase critical regime
- Both dilaton and Higgs mass loop-suppressed
- Clear prediction for direct detection
- In progress: freeze-in, leptogenesis & inflation in an extension of the model