

# **$U(1)$ -charged Dark Matter in three-Higgs-doublet models**

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**Anton Kunčinas**

Centro de Física Teórica de Partículas – CFTP and Dept de Física Instituto Superior Técnico – IST,  
Universidade de Lisboa, Portugal

In collaboration with: **P. Osland, M. N. Rebelo**  
based on [2408.02728].

**Workshop on Multi-Higgs Models**  
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Allow for spontaneous symmetry breaking, complex couplings, soft terms.

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Different symmetries may lead to the same  $V_{\mathcal{G}}$  (but likely not  $\mathcal{L}_Y$ ),  
e.g.,

$$\left. \begin{array}{l} S_3 \otimes \mathbb{Z}_2 \cong D_6 \\ [U(1) \otimes U(1)] \rtimes S_3 \\ Q_8 \ (g_1 \in \mathbb{Z}_4) \\ \{O(2), D_4, \text{CP}_2, \text{CP}_4\} \rtimes \mathbb{Z}_3 \end{array} \right\} O(2) \otimes U(1).$$

## Example of Allowed Couplings

$\mu_{11}^2$	$\mu_{22}^2$	$\mu_{33}^2$	$(\mu_{12}^2)^R$
$\lambda_{1111}$	$\lambda_{2222}$	$\lambda_{3333}$	
$\lambda_{1122}$	$\lambda_{1133}$	$\lambda_{2233}$	
$\lambda_{1221}$	$\lambda_{1331}$	$\lambda_{2332}$	

$U(1) \times U(1)$

$\left. \right\} 3+1$   
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$$[U(1) \times U(1)] +$$
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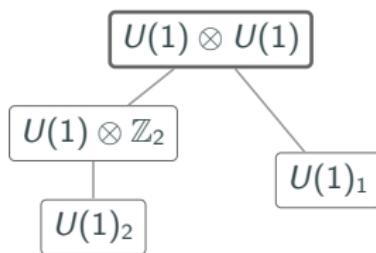
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↓

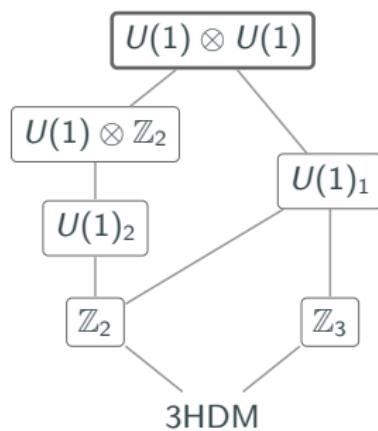
$$\begin{array}{c} [U(1) \times \mathbb{Z}_2] + \\ \boxed{\begin{array}{cc} \mu_{12}^2 & (\mu_{23}^2)^R \\ \hline \lambda_{1112} & \lambda_{1222} \\ \lambda_{1233} & \lambda_{1332} \end{array}} \\ \left. \right\} 4+1 \\ \left. \right\} 11 \end{array}$$

$= U(1)_2$

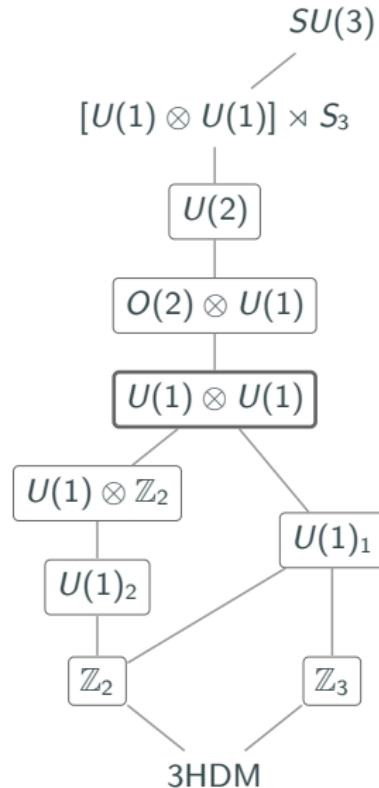
## Considered Symmetries



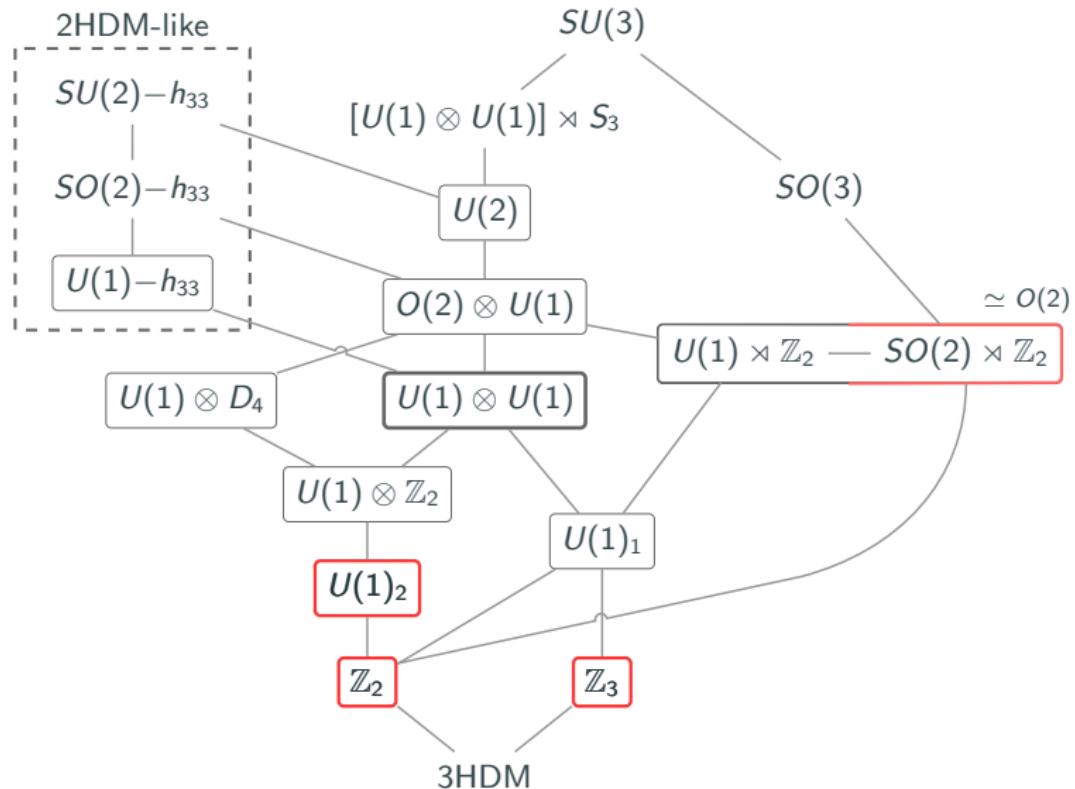
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# Implementations within $U(1) \otimes U(1)$ -3HDM

$U(1) \otimes U(1)$ -3HDM:

$$V_0 = \sum_i \mu_{ii}^2 h_{ii} + \sum_i \lambda_{i\bar{i}\bar{i}} h_{ii}^2 + \sum_{i < j} \lambda_{i\bar{i}j\bar{j}} h_{ii} h_{jj} + \sum_{i < j} \lambda_{ij\bar{j}i} h_{ij} h_{ji}, \text{ where } h_{ij} \equiv h_i^\dagger h_j.$$

Vacuum	SYM	$V$	Mixing of the neutral states	Comments
$(\hat{v}_1, \hat{v}_2, 0)$	—	$V_0$	$\{\eta_1, \eta_2\} - \{\eta_3\}$ $- \{\chi_1\} - \{\chi_2\} - \{\chi_3\}$	$m_{\eta_3} = m_{\chi_3}$ , $m_{\chi_1} = m_{\chi_2} = 0$
$(\hat{v}_1, \hat{v}_2, 0)$	—	$V_0 + \{(\mu_{12}^2)^R\}$	$\{\eta_1, \eta_2\} - \{\chi_1, \chi_2\}$ $- \{\eta_3\} - \{\chi_3\}$	$m_{\eta_3} = m_{\chi_3}$
$(v, 0, 0)$	✓	$V_0$	diagonal	$m_{\eta_2} = m_{\chi_2}$ , $m_{\eta_3} = m_{\chi_3}$

## Pitfalls: Example of $O(2)$ -3HDM

Consider two bases of  $O(2)$ :

$$V_{SO(2) \rtimes \mathbb{Z}_2} : \begin{pmatrix} \cos \theta & \sin \theta & 0 \\ \sin \theta & -\cos \theta & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \text{and} \quad V_{U(1)_1 \rtimes \mathbb{Z}_2} : \begin{pmatrix} 0 & e^{-i\theta} & 0 \\ e^{i\theta} & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}.$$

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$SO(2) \rtimes \mathbb{Z}_2 : (\hat{v}_1 e^{i\sigma}, \hat{v}_2, 0)$  and  $U(1)_1 \rtimes \mathbb{Z}_2 : (v_1, v_2, 0)$  require  
 $2\lambda_{1111} - \lambda_{1122} - \lambda_{1221} = 0$  and there are  $3 \times m_N^2 = 0$ . Symmetry increase!?

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$SO(2) \rtimes \mathbb{Z}_2 : (0, \hat{v}_2 e^{i\sigma}, \hat{v}_3) \rightarrow V_{O(2) \otimes U(1)}$ ,  $\sigma$  redundant  $\therefore$  inconsistent vacuum.

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CPV in  $SO(2) \rtimes \mathbb{Z}_2 : (0, \hat{v}_2, \hat{v}_3) + \mu_{23}^2$  (soft term) but no CPV in  $U(1)_1 \rtimes \mathbb{Z}_2$ .

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DM might be *hidden* within  $\langle h_i \rangle \neq 0$ .

# The Dark Matter Candidates

Cases *without* spontaneous symmetry breaking:

- $U(1) \otimes U(1)$        $(\nu, 0, 0)$
- $U(1)_1$                    $(\nu, 0, 0)$      $(0, 0, \nu)$
- $U(1) \otimes \mathbb{Z}_2$        $(\nu, 0, 0)$      $(0, 0, \nu)$
- $U(1)_2$                    $(\nu, 0, 0)$      $(0, 0, \nu)$      $(\nu_1, \nu_2, 0)$
- $O(2) \otimes U(1)$                    $(0, 0, \nu)$
- $O(2)$                            $(0, 0, \nu)$
- $U(1) \otimes D_4$                    $(0, 0, \nu)$
- $U(2)$                            $(0, 0, \nu)$
- $U(1)-h_{33}$                    $(\nu, 0, 0)$      $(0, 0, \nu)$
- $SO(2)-h_{33}$                    $(0, 0, \nu)$
- $SU(2)-h_{33}$                    $(0, 0, \nu)$

$[U(1) \otimes U(1)] \rtimes S_3, SO(3), SU(3)$  result in spontaneous symmetry breaking.

# The Dark Matter Candidates: Mass Degeneracies

Model	Vacuum	Symmetry	$h_{\text{SM}}$
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$$m_{\eta_3} = m_{\chi_3}$$

I-a	( $v$ , 0, 0)	$U(1) \otimes \mathbb{Z}_2$	$\eta_1$
I-b	( $v$ , 0, 0)	$U(1)_2$	$\{\eta_1, \eta_2, \chi_2\}$
I-c	( $v$ , 0, 0)	$U(1) - h_{33}$	$\eta_1$
I-d	( $\hat{v}_1, \hat{v}_2, 0$ )	$U(1)_2$	$\{\eta_1, \eta_2, \chi_1, \chi_2\}$
I-e	( $\hat{v}_1 e^{i\sigma}, \hat{v}_2, 0$ )	$U(1)_2$	$\{\eta_1, \eta_2, \chi_1, \chi_2\}$

$$m_{\eta_i} = m_{\chi_i}, \text{ for } \langle h_i \rangle = 0$$

II-a	( $v$ , 0, 0)	$U(1) \otimes U(1)$	$\eta_1$
II-b	( $v$ , 0, 0)	$U(1)_1$	$\eta_1$
II-c	(0, 0, $v$ )	$U(1) \otimes \mathbb{Z}_2$	$\eta_3$

$$m_{\{\eta_1, \eta_2\}} = m_{\{\chi_1, \chi_2\}}, \eta_i \text{ and } \chi_i \text{ mix separately}$$

II-d	(0, 0, $v$ )	$U(1)_1$	$\eta_3$
II-e	(0, 0, $v$ )	$O(2)$	$\eta_3$

$$m_{\{\eta_1, \eta_2, \chi_1, \chi_2\}}, \text{ all states mix together}$$

II-f	(0, 0, $v$ )	$U(1)_2$	$\eta_3$
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$$m_{\eta_1} = m_{\eta_2} = m_{\chi_1} = m_{\chi_2}$$

II-g	(0, 0, $v$ )	$O(2) \otimes U(1)$	$\eta_3$
II-h	(0, 0, $v$ )	$U(1) \otimes D_4$	$\eta_3$
II-i	(0, 0, $v$ )	$U(2)$	$\eta_3$

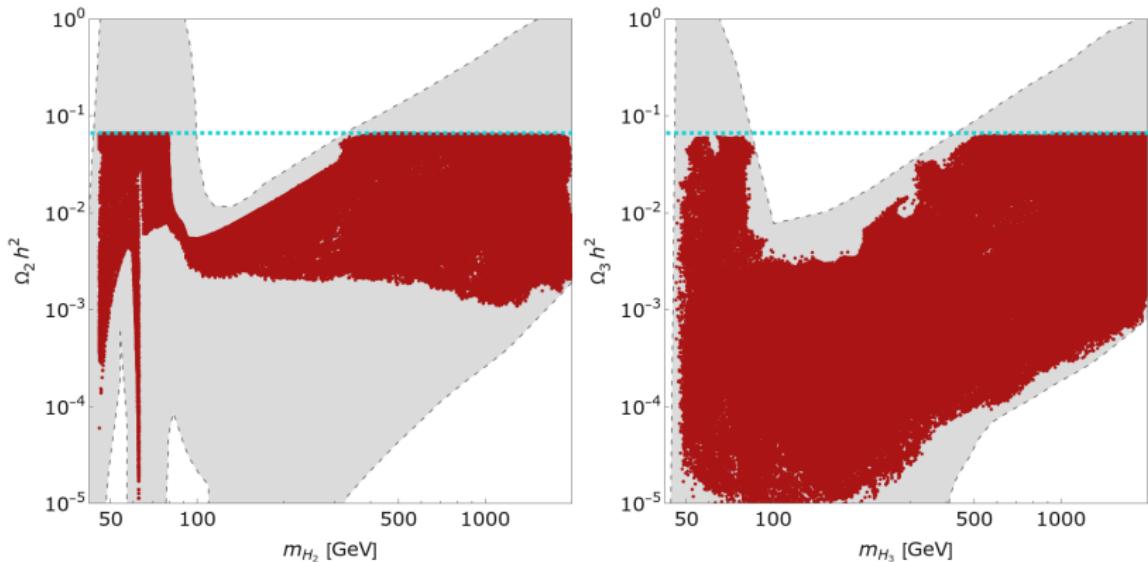
[See P. Osland's talk.]

# The Dark Matter Candidates: Scalar Couplings

Symmetry	$(\nu, 0, 0)$	$(0, 0, \nu)$	$(\nu_1, \nu_2, 0)$
$U(1) \otimes U(1)$	$7 + 32 (5 + 10)$		
$U(1)_1$	$15 + 40 (8 + 13)$	$9 + 53 (6 + 24)$	
$U(1) \otimes \mathbb{Z}_2$	$7 + 32 (6 + 11)$	$7 + 38 (5 + 16)$	
$U(1)_2$	$24 + 55 (20 + 40)$	$12 + 73 (8 + 36)$	$\mathbb{R} : 16 + 41 (13 + 24)$ $\mathbb{C} : 24 + 55 (20 + 40)$
$O(2) \otimes U(1)$		$7 + 32 (3 + 7)$	
$O(2)$		$7 + 39 (4 + 12)$	
$U(1) \otimes D_4$		$7 + 34 (3 + 11)$	
$U(2)$		$7 + 32 (3 + 6)$	
$U(1) - h_{33}$	$4 + 16 (3 + 5)$	$1 + 26 (1 + 6)$	
$SO(2) - h_{33}$		$1 + 26 (1 + 5)$	
$SU(2) - h_{33}$		$1 + 26 (1 + 4)$	

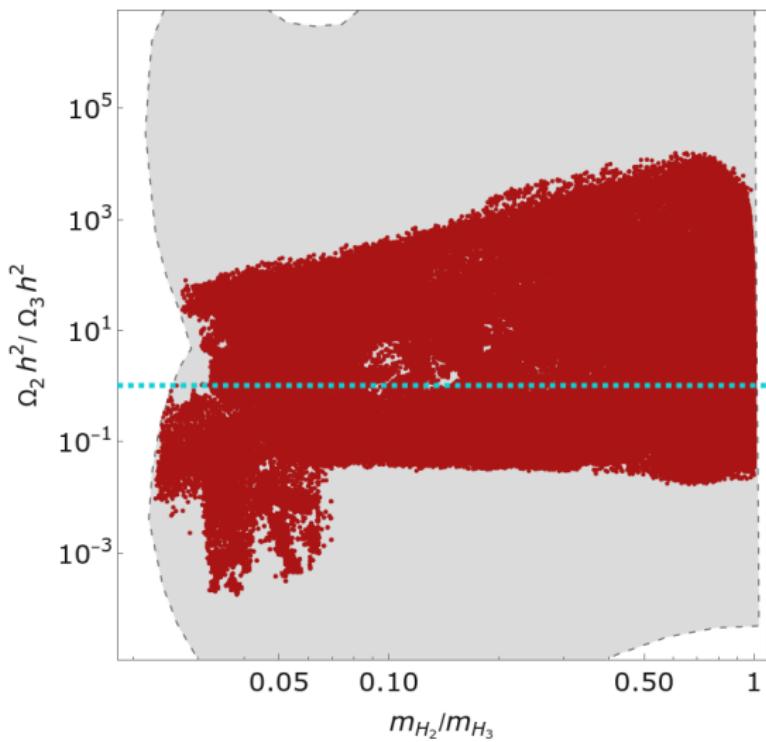
# Dark Matter in $U(1) \otimes U(1)$ -3HDM

There are 4 (2) DM candidates,  $\Omega_{\text{DM}} h^2 = \Omega_2 h^2 + \Omega_3 h^2$ .



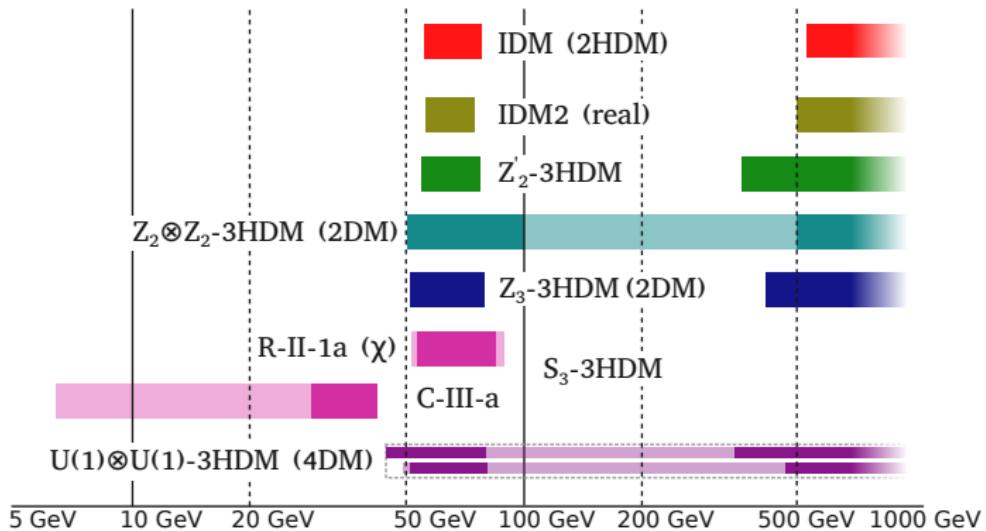
- No points when  $m_{H_2} \in [100; 300]$  GeV and  $m_{H_3} \in [100; 450]$  GeV;
- In  $[50; 80] \cup [300; 2000]$  GeV comparable  $m_{H_i}^2$  and  $\Omega_i h^2$ ;
- Regions where  $\Omega h^2$  is dominated by a single mass scale,  
 $m_{H_2} = [45; 80] \cup [330; 2000]$  GeV and  $m_{H_3} = [52; 80] \cup [470; 2000]$  GeV;

# Dark Matter in $U(1) \otimes U(1)$ -3HDM



# Updated Stock of Dark Matter in 3HDMs

## SCALAR DM MASS RANGES



IDM: (A. Belyaev, G. Cacciapaglia, I. P. Ivanov, F. Rojas-Abatte, M. Thomas, 2016);  
(J. Kalinowski, W. Kotlarski, T. Robens, D. Sokolowska, A. F. Zarnecki, 2018);

IDM2: (M. Merchant, M. Sher, 2019);

$\mathbb{Z}_n$ -3HDM: (A. Cordero-Cid, J. Hernández-Sánchez, V. Keus, S. F. King, S. Moretti, D. Rojas, D. Sokołowska, 2014-2022);

$\mathbb{Z}_2 \otimes \mathbb{Z}_2$ -3HDM: (R. Boto, P. N. Figueiredo, J. C. Romão, J. Silva, 2024);

$S_3$ -3HDM: (A. Kunčinas, O. M. Ógreid, P. Osland, M. N. Rebelo, 2021-2023);

- We classified and identified models, embedding  $U(1)$ -stabilised DM in 3HDMs. These models contain mass-degenerate pairs of DM candidates, due to the unbroken  $U(1)$ . Such classification and identification of models is useful for model builders interested in three-Higgs-doublet models stabilised by continuous symmetries;
- We performed a numerical scan of the  $U(1) \otimes U(1)$ -3HDM. Within the model there is a multi-component DM sector, with two different mass scales. We found that there are possible solutions throughout a broad DM mass range, 45–2000 GeV;

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## Appendix: on $\text{CP}_2 \rightarrow D_4$

$\text{CP}_2$ :  $\{h_1 \rightarrow h_2^*, h_2 \rightarrow -h_1^*, h_3 \rightarrow h_3^*\}$ .

$$\begin{aligned} V_{\text{CP}_2} = & \mu_{11}^2(h_{11} + h_{22}) + \mu_{33}^2 h_{33} + \lambda_{1111}(h_{11}^2 + h_{22}^2) + \lambda_{3333} h_{33}^2 \\ & + \lambda_{1122} h_{11} h_{22} + \lambda_{1133}(h_{11} h_{33} + h_{22} h_{33}) + \lambda_{1221} h_{12} h_{21} \\ & + \lambda_{1331}(h_{13} h_{31} + h_{23} h_{32}) + \lambda_{1212}(h_{12}^2 + h_{21}^2) + \left\{ \lambda_{1313}(h_{13}^2 + h_{23}^2) + \text{h.c.} \right\} \\ & + \lambda_{1112}(h_{11} h_{12} + h_{11} h_{21} - h_{12} h_{22} - h_{21} h_{22}). \end{aligned}$$

Perform a basis transformation  $\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} i & 1 & 0 \\ 1 & i & 0 \\ 0 & 0 & \sqrt{2} \end{pmatrix} \begin{pmatrix} h'_1 \\ h'_2 \\ h'_3 \end{pmatrix}$ .

$$\begin{aligned} V_{D_4} = & \mu_{11}^2(h_{11} + h_{22}) + \mu_{33}^2 h_{33} + \lambda_{1111}(h_{11}^2 + h_{22}^2) + \lambda_{3333} h_{33}^2 + \lambda_{1122} h_{11} h_{22} \\ & + \lambda_{1133}(h_{11} h_{33} + h_{22} h_{33}) + \lambda_{1221} h_{12} h_{21} + \lambda_{1331}(h_{13} h_{31} + h_{23} h_{32}) \\ & + \lambda_{1212}(h_{12}^2 + h_{21}^2) + \left\{ \lambda_{1313}(h_{13}^2 + h_{23}^2) + \text{h.c.} \right\}. \end{aligned}$$

$\text{CP}_2$ -symmetric 3HDM has a redundant quartic coupling  $\lambda_{1112}$ .

## Appendix: $O(2) \otimes U(1)$

Invariant under:

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \begin{pmatrix} 0 & e^{i\theta_1} & 0 \\ e^{i\theta_2} & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} h'_1 \\ h'_2 \\ h'_3 \end{pmatrix}.$$

$$\begin{aligned} V = & \mu_{11}^2(h_{11} + h_{22}) + \mu_{33}^2 h_{33} + \lambda_{1111}(h_{11}^2 + h_{22}^2) + \lambda_{3333} h_{33}^3 + \lambda_{1122} h_{11} h_{22} \\ & + \lambda_{1133}(h_{11} h_{33} + h_{22} h_{33}) + \lambda_{1221} h_{12} h_{21} + \lambda_{1331}(h_{13} h_{31} + h_{23} h_{32}). \end{aligned}$$

$S_3 \otimes \mathbb{Z}_2 \simeq D_6$ . However,  $D_n$ , with  $n \geq 5$  result in identical  $V$ .

Perform a basis change:

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} -e^{-i\frac{\pi}{4}} & -e^{i\frac{\pi}{4}} & 0 \\ e^{-i\frac{\pi}{4}} & -e^{i\frac{\pi}{4}} & 0 \\ 0 & 0 & \sqrt{2} \end{pmatrix} \begin{pmatrix} h'_1 \\ h'_2 \\ h'_3 \end{pmatrix}.$$

$$\begin{aligned} V = & \mu_{11}^2(h_{11} + h_{22}) + \mu_{33}^2 h_{33} + \lambda_a(h_{11}^2 + h_{22}^2) + \lambda_{3333} h_{33}^3 + \lambda_b h_{12} h_{21} \\ & + \lambda_{1331}(h_{13} h_{31} + h_{23} h_{32}) + \lambda_c h_{11} h_{22} + \lambda_{1133}(h_{11} h_{33} + h_{22} h_{33}) - \frac{1}{2}\Lambda(h_{12}^2 + h_{21}^2). \end{aligned}$$

Invariance under  $g = \begin{pmatrix} 0 & e^{i\alpha} & 0 \\ -e^{-i\alpha} & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$ .

## Appendix: $[U(1) \otimes U(1)] \rtimes S_3$

Invariant under:

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \begin{pmatrix} 0 & e^{i\theta_1} & 0 \\ 0 & 0 & e^{i\theta_2} \\ 1 & 0 & 0 \end{pmatrix} \begin{pmatrix} h'_1 \\ h'_2 \\ h'_3 \end{pmatrix}.$$

$$\begin{aligned} V &= \mu_{11}^2(h_{11} + h_{22} + h_{33}) + \lambda_{1111}(h_{11}^2 + h_{22}^2 + h_{33}^2) \\ &\quad + \lambda_{1122}(h_{11}h_{22} + h_{22}h_{33} + h_{33}h_{11}) + \lambda_{1221}(h_{12}h_{21} + h_{23}h_{32} + h_{31}h_{13}) \\ &= \mu_{11}^2 \sum_i h_{ii} + \lambda_{1111} \sum_i h_{ii}^2 + \lambda_{1122} \sum_{i < j} h_{ii}h_{jj} + \lambda_{1221} \sum_{i < j} h_{ij}h_{ji}. \end{aligned}$$

$$g = \begin{pmatrix} 0 & e^{i\theta_1} & 0 \\ 0 & 0 & e^{i\theta_2} \\ e^{-i(\theta_1+\theta_2)} & 0 & 0 \end{pmatrix}.$$

## Appendix: $U(1) \otimes D_4$

$$V_{U(1) \otimes Z_2}^{\text{ph}} = \lambda_{1212} h_{12}^2 + \text{h.c.},$$

$$\begin{aligned} V = & \mu_{11}^2(h_{11} + h_{22}) + \mu_{33}^2 h_{33} + \lambda_{1111}(h_{11}^2 + h_{22}^2) + \lambda_{3333} h_{33}^2 + \lambda_{1122} h_{11} h_{22} \\ & + \lambda_{1133}(h_{11} h_{33} + h_{22} h_{33}) + \lambda_{1221} h_{12} h_{21} + \lambda_{1331}(h_{13} h_{31} + h_{23} h_{32}) \\ & + \lambda_{1212}(h_{12}^2 + h_{21}^2). \end{aligned}$$

Invariance under  $\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \begin{pmatrix} 0 & -e^{i\alpha} & 0 \\ e^{i\alpha} & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} h'_1 \\ h'_2 \\ h'_3 \end{pmatrix}$

suggests that  $\{\lambda_{1112}(h_{11} h_{12} - h_{21} h_{22}) + \text{h.c.}\}$  is allowed. Change of basis

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} e^{\frac{i\pi}{4}} & e^{\frac{i\pi}{4}} & 0 \\ -e^{-\frac{i\pi}{4}} & e^{-\frac{i\pi}{4}} & 0 \\ 0 & 0 & \sqrt{2} \end{pmatrix} \begin{pmatrix} h'_1 \\ h'_2 \\ h'_3 \end{pmatrix} \text{ yields } \lambda_{1212} \in \mathbb{C}.$$

As pointed out by I. Ivanov, this should better be denoted as a quotient group.

## Appendix: Different Models

Underlying symmetry	Reference	Indep. couplings	Allowed couplings and necessary relations
$\mathbb{Z}_2$	Table 16	4 + 17	$V_0, \mu_{12}^2, \lambda_{ijij}, \lambda_{1112}, \lambda_{1222}, \lambda_{1233}, \lambda_{1323}, \lambda_{1332}$
$U(1)_2$	Table 4	4 + 14	$V_0, \mu_{12}^2, \lambda_{1212}, \lambda_{1112}, \lambda_{1222}, \lambda_{1233}, \lambda_{1332}$
$\mathbb{Z}_3$	Table 17	3 + 12	$V_0, \lambda_{1323}, \lambda_{1213}, \lambda_{1232}$
$U(1)_1$	Table 2	3 + 10	$V_0, \lambda_{1323}$
$U(1) \otimes \mathbb{Z}_2$	Table 3	3 + 10	$V_0, \lambda_{1212}$
$U(1) \otimes U(1)$	Table 1	3 + 9	$V_0$
$O(2)_{[SO(2) \rtimes \mathbb{Z}_2]}$	Table 8	2 + 7	$V_{O(2) \otimes U(1)}, \lambda_{1212} = \Lambda, \lambda_{1313} = \lambda_{2323}$
$O(2)_{[U(1) \rtimes \mathbb{Z}_2]}$	Table 9	2 + 7	$V_{O(2) \otimes U(1)}, \lambda_{1323}$
$U(1) \otimes D_4$	Table 10	2 + 7	$V_{O(2) \otimes U(1)}, \lambda_{1212}$
$O(2)_{[SO(2) \rtimes \mathbb{Z}_2]} \otimes U(1)$	Table 5	2 + 6	$V_{O(2) \otimes U(1)}, \lambda_{1212} = \Lambda$
$O(2)_{[U(1) \rtimes \mathbb{Z}_2]} \otimes U(1)$	Table 6	2 + 6	$V_{O(2) \otimes U(1)}$
$U(2)$	Table 11	2 + 5	$\mu_{11}^2 = \mu_{22}^2, \mu_{33}^2, \lambda_{1111} = \lambda_{2222}, \lambda_{3333}, \Lambda = 0, \lambda_{1133} = \lambda_{2233}, \lambda_{1221} \\ \lambda_{1331} = \lambda_{2332}$
$[U(1) \otimes U(1)] \rtimes S_3$	Table 7	1 + 3	$\mu_{11}^2 = \mu_{22}^2 = \mu_{33}^2, \lambda_{1111} = \lambda_{2222} = \lambda_{3333}, \lambda_{1122} = \lambda_{1133} = \lambda_{2233}, \\ \lambda_{1221} = \lambda_{1331} = \lambda_{2332}$
$SO(3)$	Table 12	1 + 3	$\mu_{11}^2 = \mu_{22}^2 = \mu_{33}^2, \lambda_{1111} = \lambda_{2222} = \lambda_{3333}, \lambda_{1122} = \lambda_{1133} = \lambda_{2233}, \\ \lambda_{1221} = \lambda_{1331} = \lambda_{2332}, \lambda_{ijij} = \Lambda$
$SU(3)$	Table 13	1 + 2	$\mu_{11}^2 = \mu_{22}^2 = \mu_{33}^2, \lambda_{1111} = \lambda_{2222} = \lambda_{3333}, \lambda_{1122} = \lambda_{1133} = \lambda_{2233}, \\ \lambda_{ijji} = 2\lambda_{1111} - \lambda_{1122}$

## Appendix: Dark Matter in $U(1) \otimes U(1)$ -3HDM

$$V_{\text{IDM}} = \mu_{11}^2 h_{11} + \mu_{22}^2 h_{22} + \lambda_{1111} h_{11}^2 + \lambda_{2222} h_{22}^2 + \lambda_{1122} h_{11} h_{22} + \lambda_{1221} h_{12} h_{21} \\ + \lambda_{1212} (h_{12}^2 + h_{21}^2).$$

$$V_{U(1) \otimes U(1)} = \frac{1}{2} \mu_{11}^2 h_{11} + \mu_{22}^2 h_{22} + \frac{1}{2} \lambda_{1111} h_{11}^2 + \lambda_{2222} h_{22}^2 + \lambda_{1122} h_{11} h_{22} + \lambda_{1221} h_{12} h_{21} \\ + \frac{1}{2} \mu_{11}^2 h_{11} + \mu_{33}^2 h_{33} + \frac{1}{2} \lambda_{1111} h_{11}^2 + \lambda_{3333} h_{33}^2 + \lambda_{1133} h_{11} h_{33} + \lambda_{1331} h_{13} h_{31} \\ + \lambda_{2233} h_{22} h_{33} + \lambda_{2332} h_{23} h_{32}.$$

Masses:

$$m_{h_i^+}^2 = \mu_{ii}^2 + \frac{1}{2} \lambda_{11ii} v^2, \text{ for } i = \{2, 3\},$$

$$m_{\eta_i}^2 = m_{\chi_i}^2 \equiv m_{H_i}^2 = \mu_{ii}^2 + \frac{1}{2} (\lambda_{11ii} + \lambda_{1ii1}) v^2, \text{ for } i = \{2, 3\}.$$

$Z$  interactions:  $\mathcal{L}_{VHH} = -\frac{g}{2 \cos \theta_W} Z^\mu \eta_i \overleftrightarrow{\partial}_\mu \chi_i.$

Higgs portal:  $g(\eta_i^2 h) = g(\chi_i^2 h) = v (\lambda_{11ii} + \lambda_{1ii1}).$

## Appendix: $U(1) \otimes U(1)$ Scan

- All of the additional scalars can be as heavy as 2 TeV;
- The DM candidates are associated with the neutral states of the  $h_2$  and  $h_3$  doublets,  $m_{H_2} \leq m_{H_3}$ ;
- LEP constraints. We allow for  $m_{H_i^\pm} \geq 70$  GeV and assume  $m_{H_i} > \frac{1}{2}m_Z$ ;
- Perturbativity, Unitarity ( $8\pi$ ), Stability conditions;
- Electroweak precision observables  $S$  and  $T$ ;
- $\text{Br}(h \rightarrow \text{invisible}) \leq 0.1$ ;
- LHC searches implemented in HiggsTools;
- DM relic density is evaluated as  $\Omega_{\text{DM}} h^2 = 2\Omega_2 h^2 + 2\Omega_3 h^2$ ;
- Direct DM searches based on XENONnT and LUX-ZEPLIN;
- Indirect DM searches.

## Appendix: Mass Scatter Plots of $U(1) \otimes U(1)$ -3HDM

